

# Superconducting Transition Edge Sensors (TES)

By:

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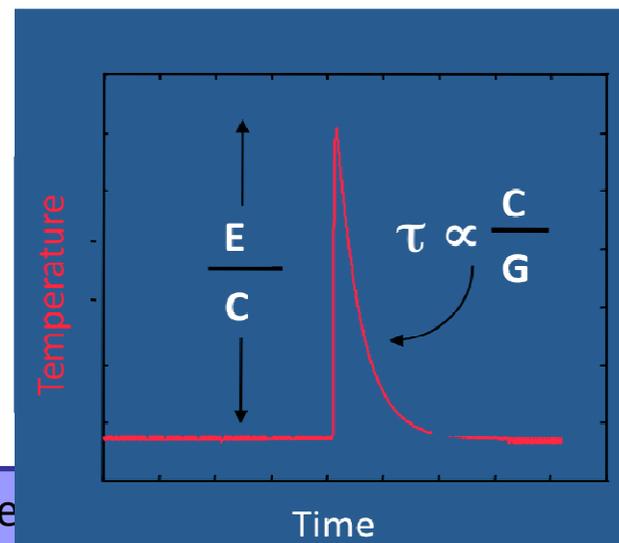
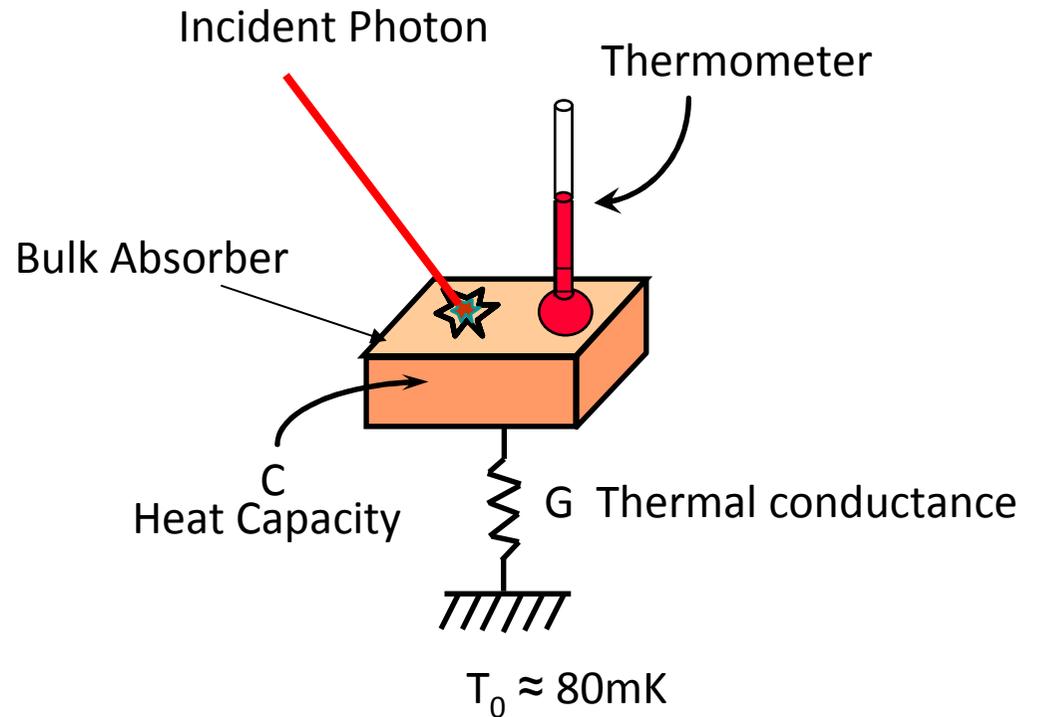
## Outline

- **Introduction**
- **Applications:**
  - **Detectors for Astronomy**
  - **Microcalorimeters for Gamma-ray Spectroscopy**
- **Optics Alignment Tool**
- **Conclusions**

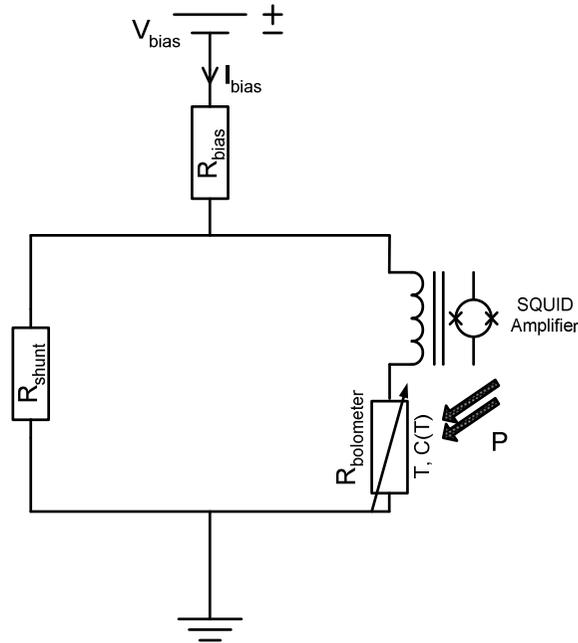
- Thermometer and absorber are connected by a weak thermal link to a heat sink.

- Incoming energy is converted to heat in the absorber:  
 $\Delta T = T - T_0 = E/C.$

- Temperature rise decays as power in absorber flows out to the heat sink  
 $\tau = C/G.$



## Voltage biased circuit: TES:

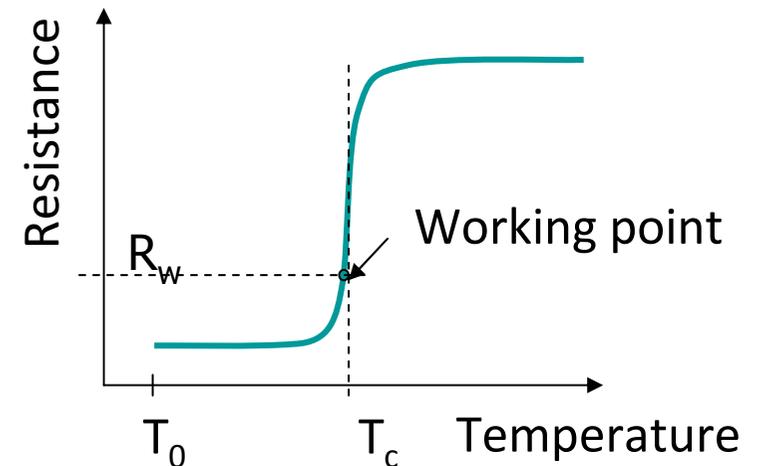


- ➡ The strong negative ETF maintains the transition temperature.
- ➡ Sharp transition speeds up the detector so lower time constant.
- ➡ Higher sensitivity.
- ➡ SQUIDs and multiplexing of SQUID electronics.

### Strong Negative Electrothermal Feedback

1)  $P_{\text{Total}} = P_{\text{Signal}} + P_{\text{Bias}}$  where  $P_{\text{Bias}} = V_0^2/R$   
 $T \uparrow \Rightarrow R \uparrow \Rightarrow P_{\text{Bias}} \downarrow \Rightarrow P_{\text{Total}} \downarrow \Rightarrow T \downarrow$

2) Faster bolometer response  $\tau_0 = \frac{C}{G}$   $\tau = \frac{\tau_0}{L+1}$



## Detectors:

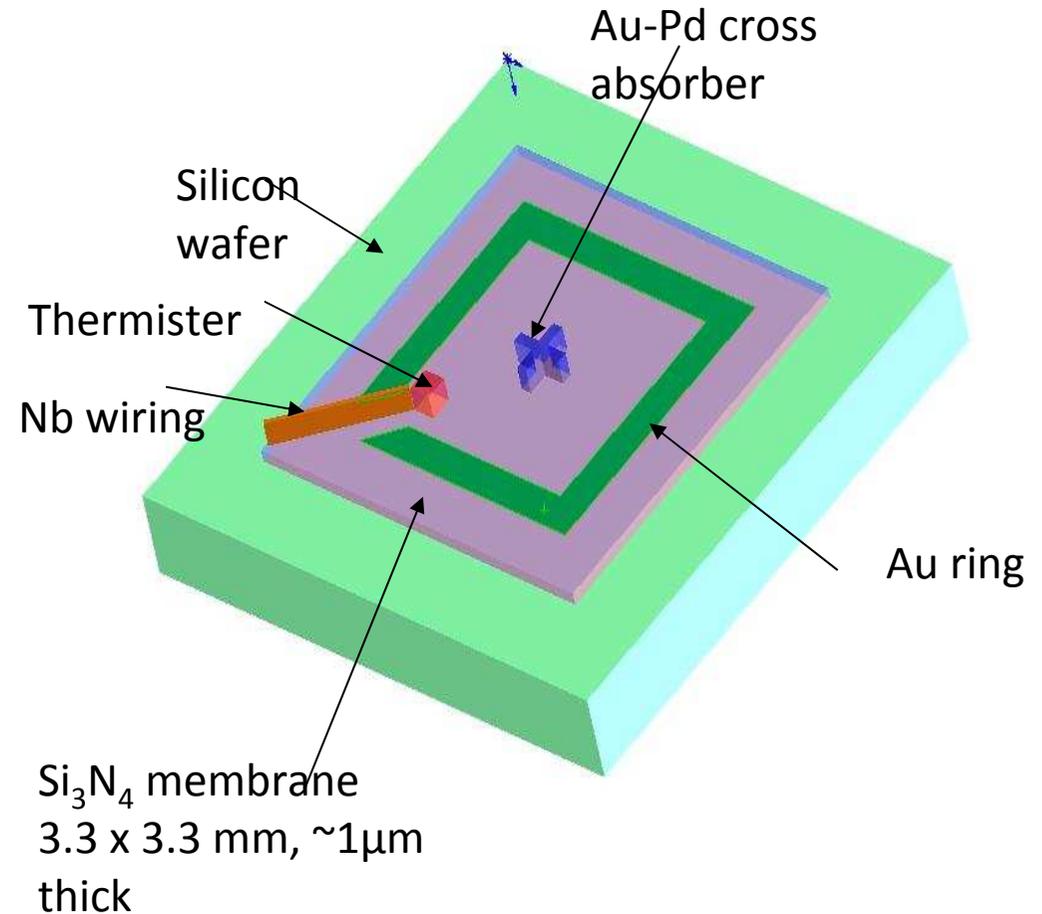
Thin film thermometers.

Bi-layer thermometers.

MPIfR → Gold-Palladium (8 nm)/Molybdenum (60 nm) →  $450 \pm 25$  mK

NIST → Copper (0.1 μm)/ Molybdenum (0.2 μm) →  $120 \pm 10$  mK

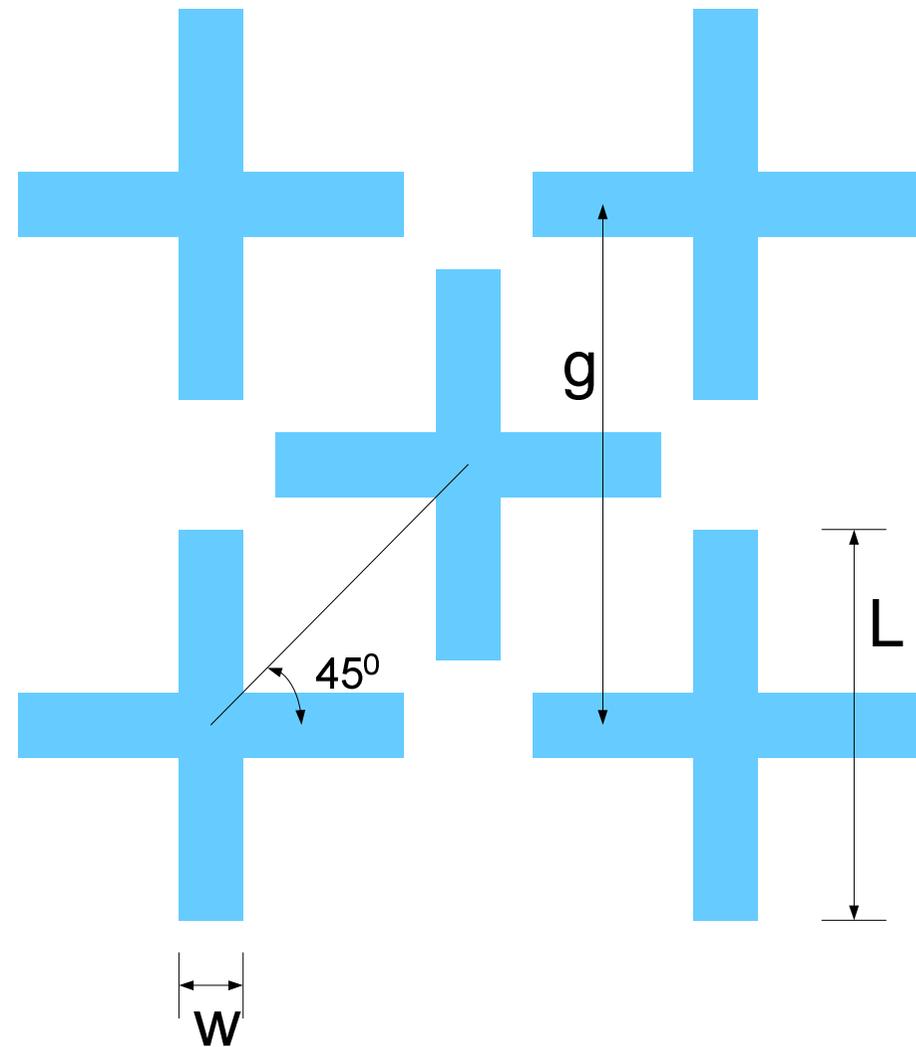
- Au-Pd / Mo bilayer structure thermistor
- Transition temperature  $T_c \approx 450$  mK
- Thermistor is mounted on silicon nitride ( $\text{Si}_3\text{N}_4$ ) membrane.



## Au-Pd cross absorber

- Gold-palladium alloy
- Surface resistance =  $10 \Omega/\square$

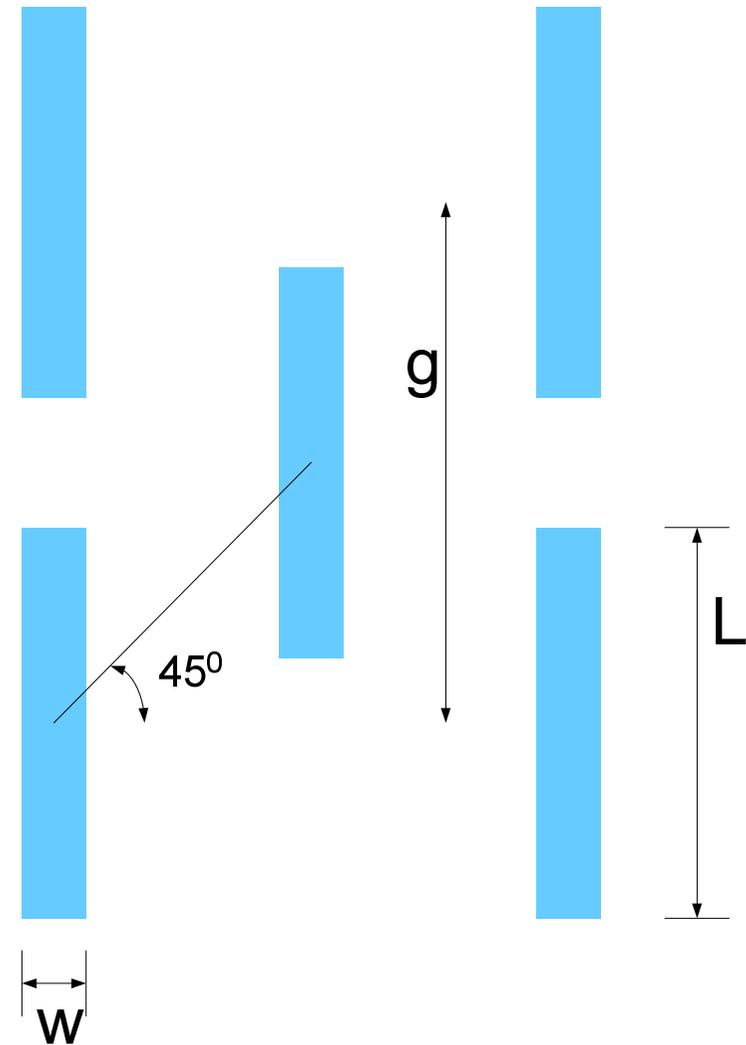
$$A = \left| \frac{4n}{(1+n)^2} \right| \frac{4Z_0 n R}{|R(1+n) + Z_0|^2}$$



## Au-Pd cross absorber

- Gold-palladium alloy
- Surface resistance =  $10 \Omega/\square$

Polarization sensitive absorbers

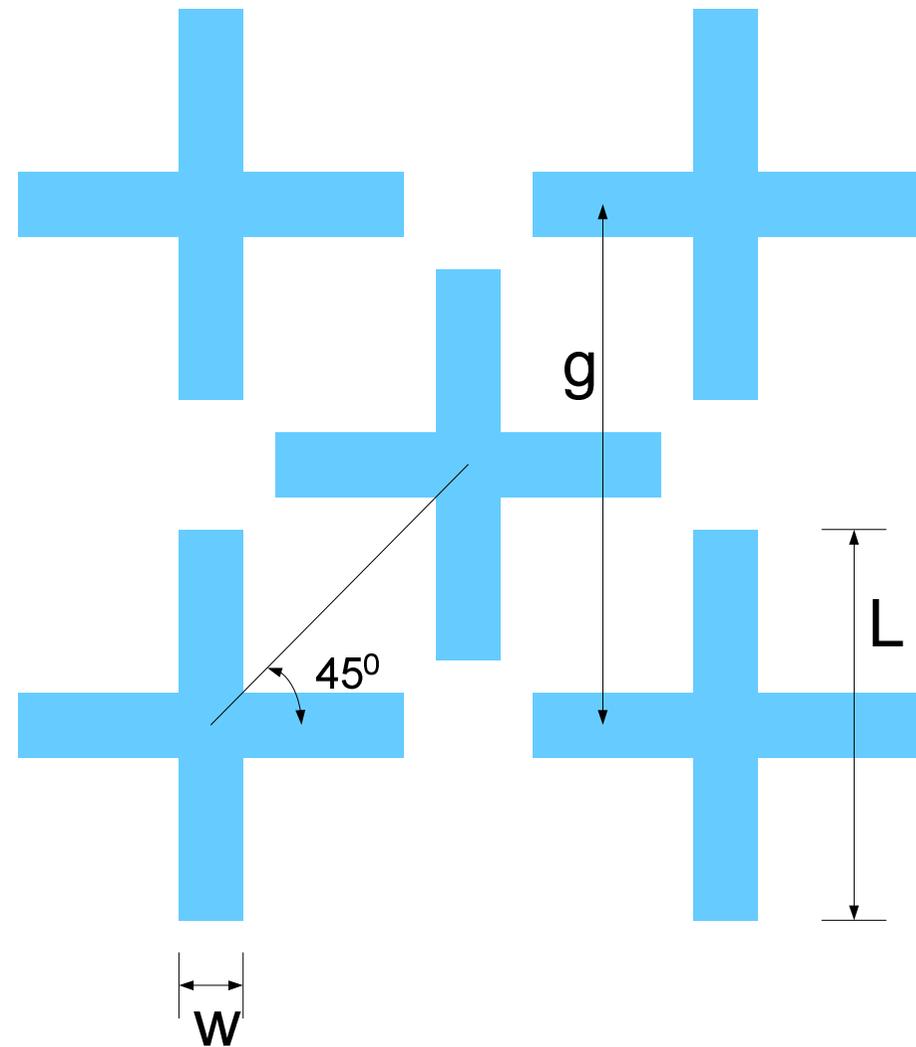


## Au-Pd cross absorber

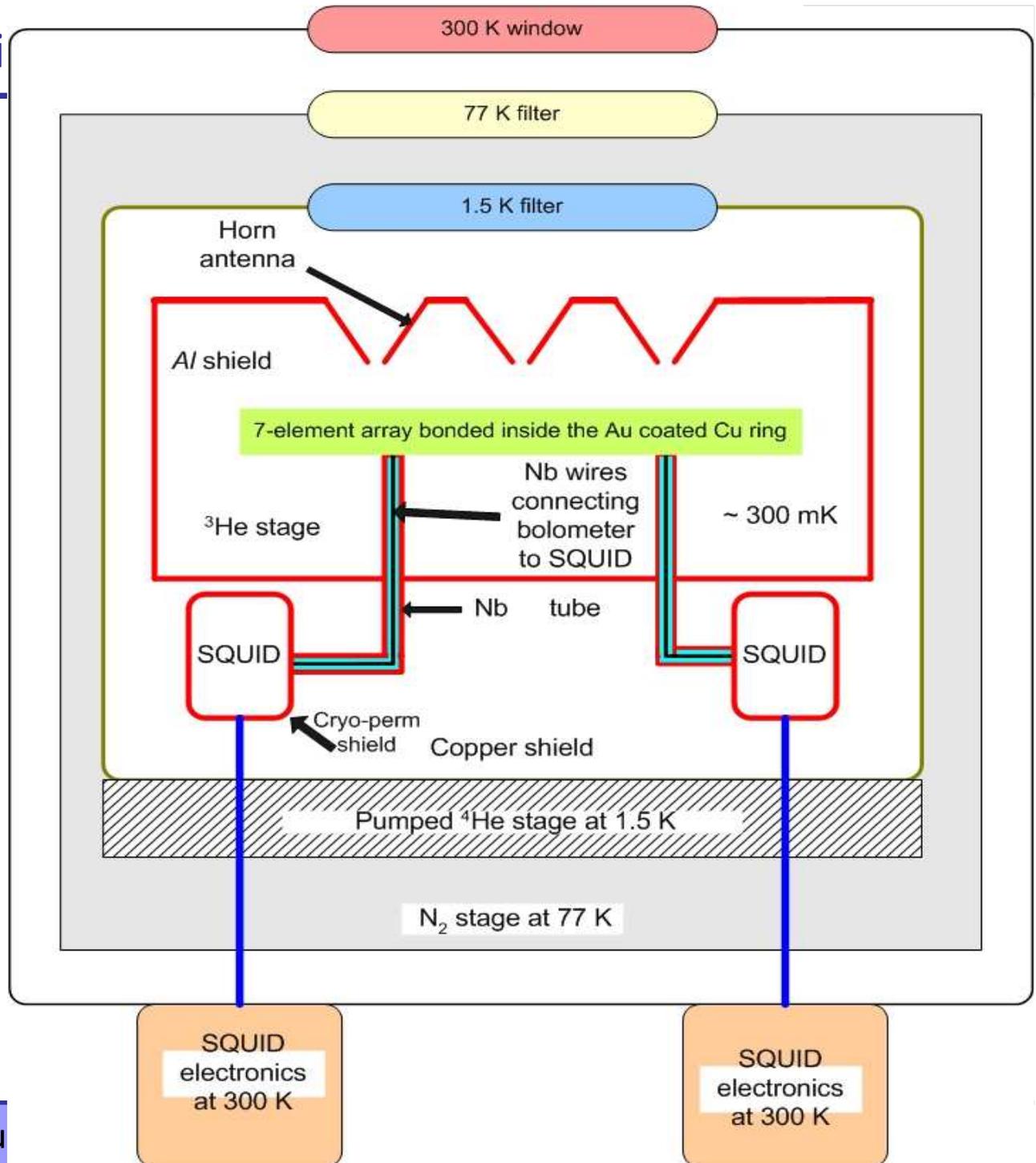
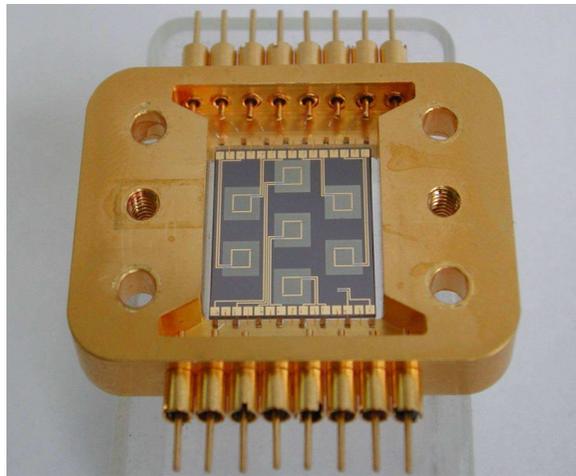
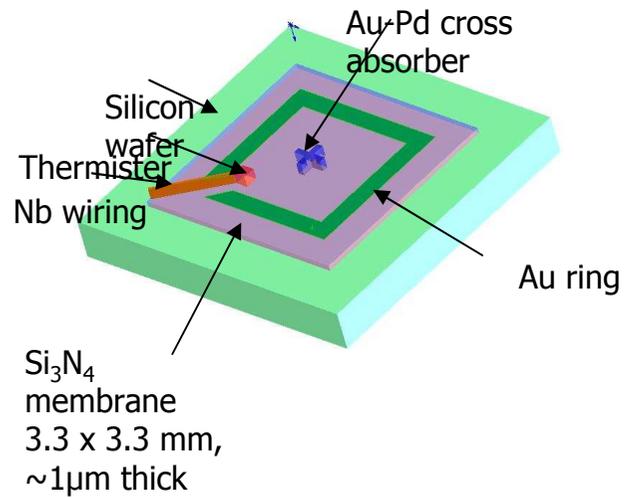
- Gold-palladium alloy
- Surface resistance =  $10 \Omega/\square$

### Advantages:

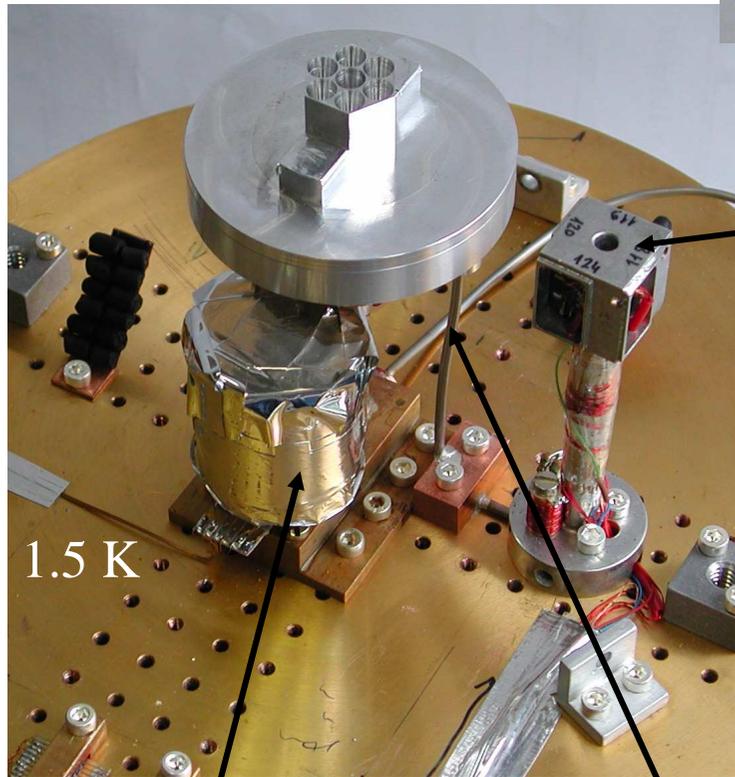
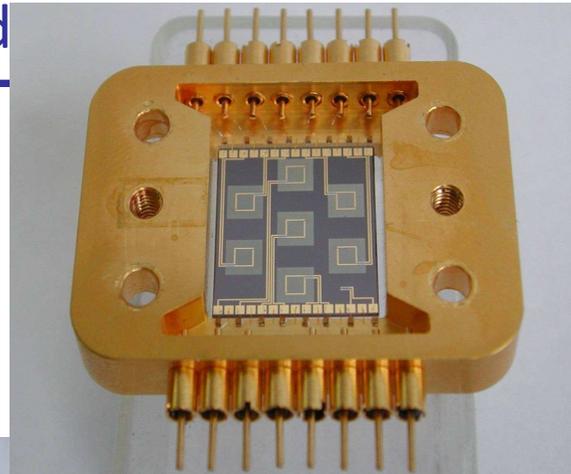
- Frequency selective absorber,  $L = \lambda/2$
- Polarization sensitive absorber



# Application: Astronomy



Superconducting aluminium  
Hornarray/cavity at 0.3 K



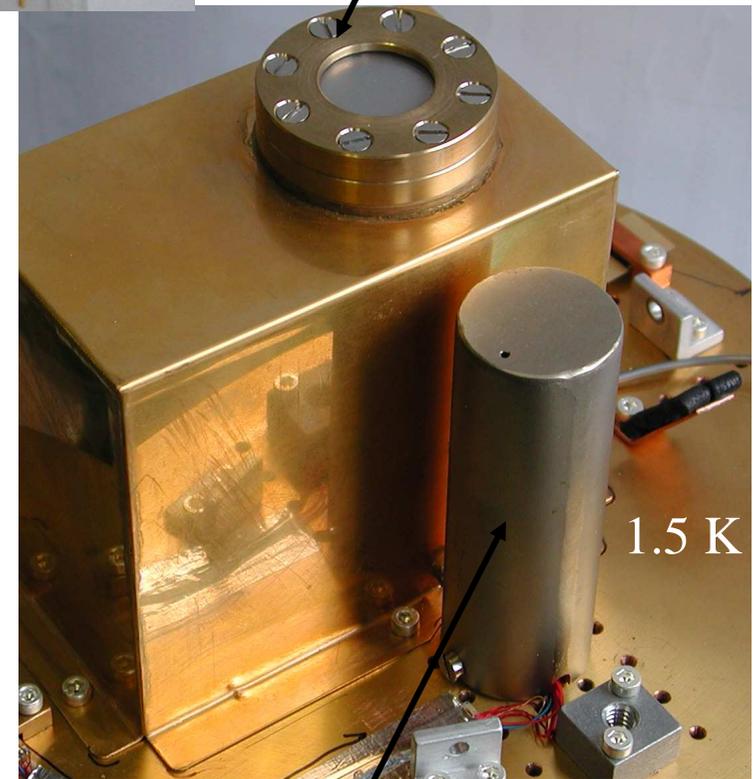
4 SQUIDs  
at 1.5 K

1.5 K

Sorption pump

Niobium tube

1.5 K filter



1.5 K

Cryo-perm shield

## TES equations:

$$S_i \equiv \frac{\delta I}{\delta P} = -\frac{1}{V_{BIAS}} \cdot \frac{L}{(L+1)} \cdot \frac{1}{(1+i\omega\tau)} \quad \text{current response}$$

$$L(\omega) = \frac{P_{BIAS} \cdot \alpha}{G(1+i\omega\tau_0)} \quad \text{AC open loop gain}$$

- 1) Structuring of  $\text{Si}_3\text{N}_4$  membrane.
- 2) Estimation of  $G$ .

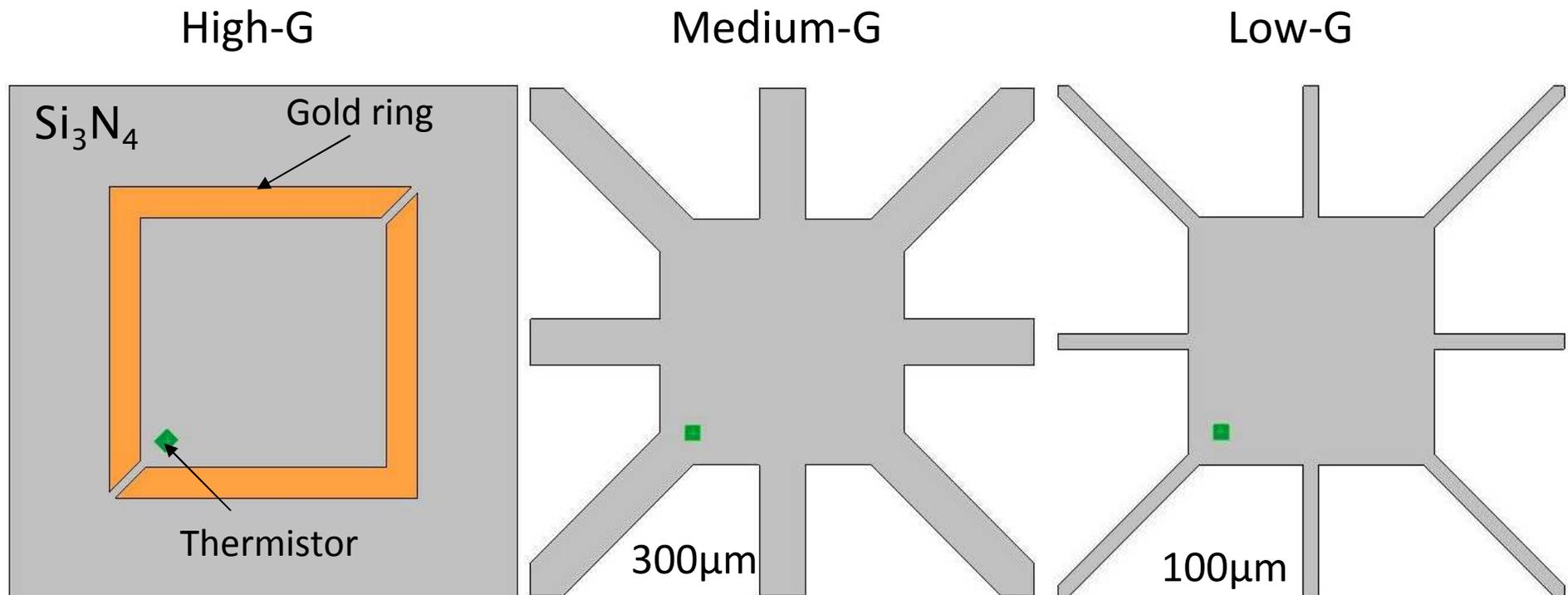
Finite Element Analysis

(WWW.COSMOSM.COM)

## Silicon nitride membrane

- The thermal conductance is tuned by structuring the silicon nitride membrane.

$$\text{sensitivity} \propto \frac{1}{\text{thermal conductance}}$$

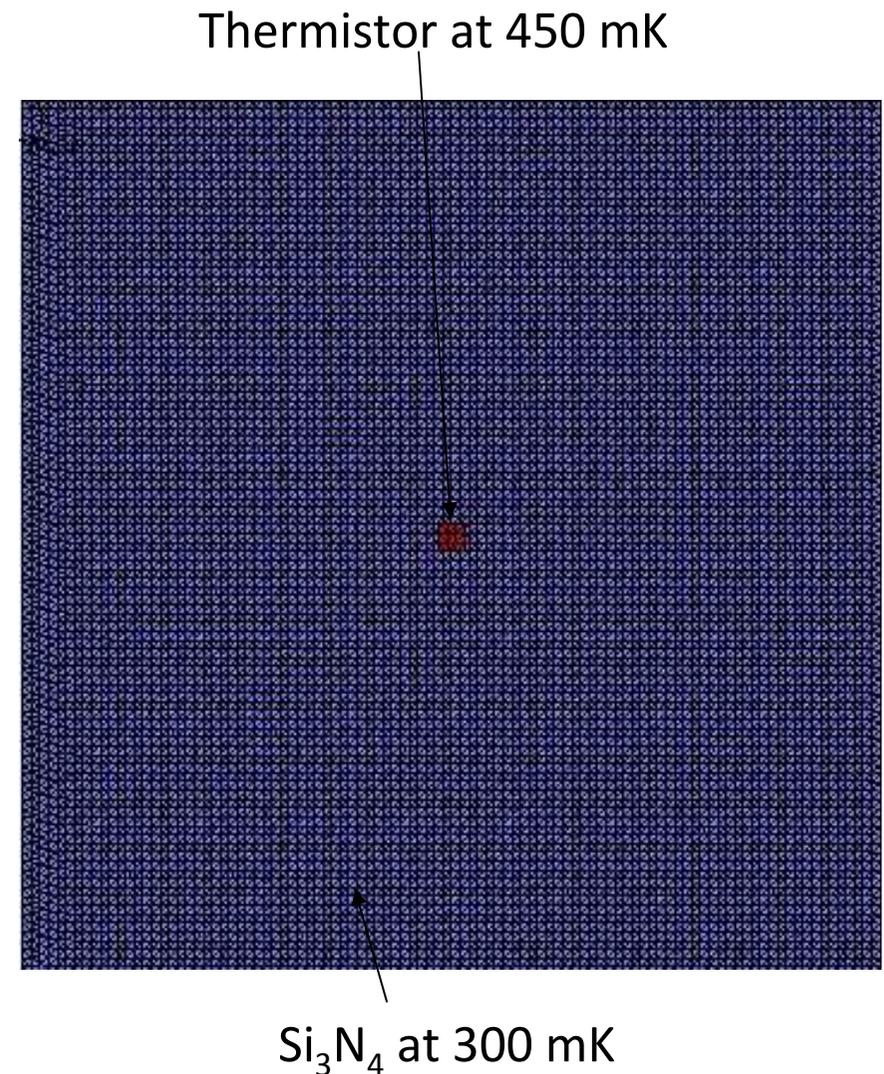


## Thermal modeling

- Thermal analysis using COSMOS

$$G_{so} = k(A/L)$$

- Thermal conductivity ( $k$ ) values are obtained from Holmes et al.  
Applied Physics letters 72, 18, 1998

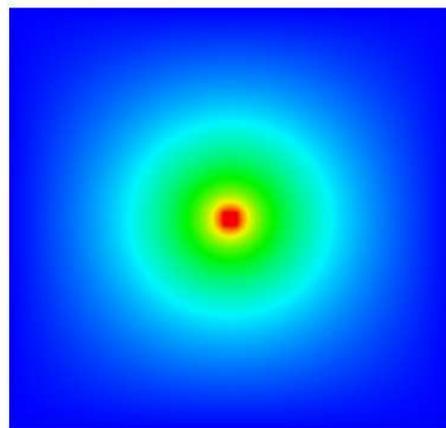
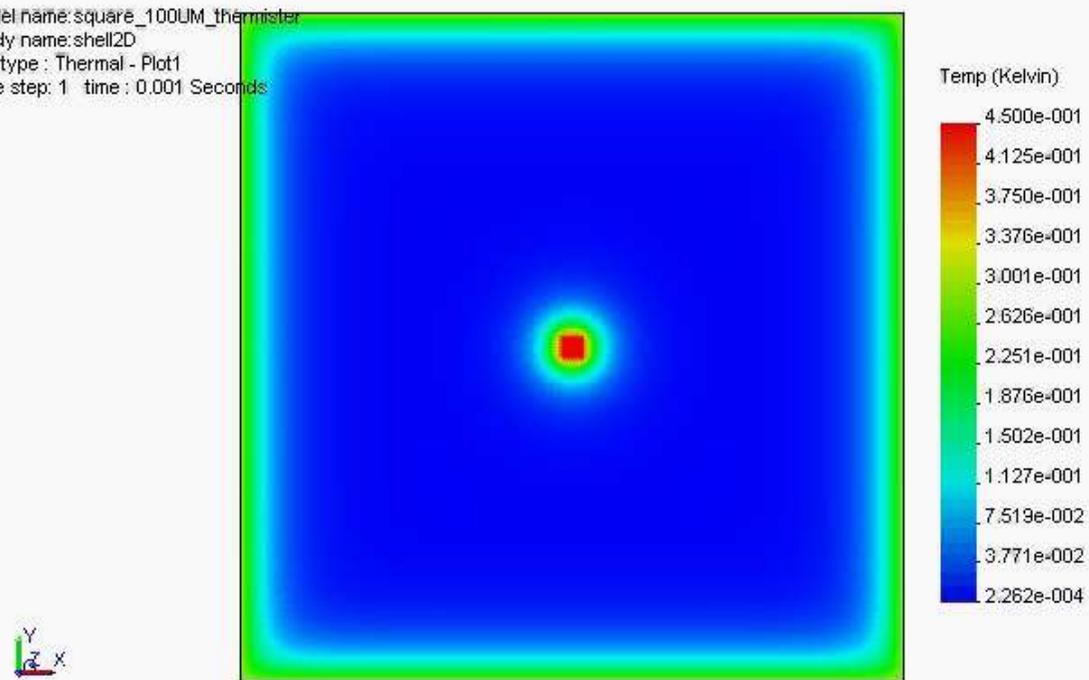


## Thermal modeling

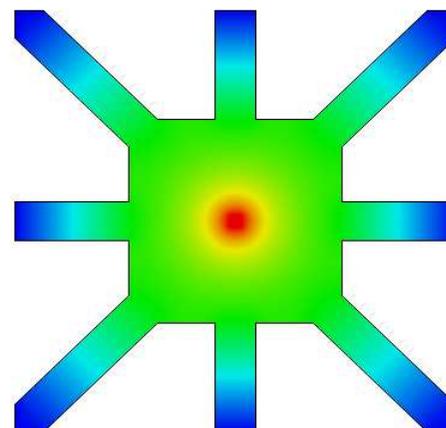
-COSMOS

Bolometer response  
to the periodic signal.

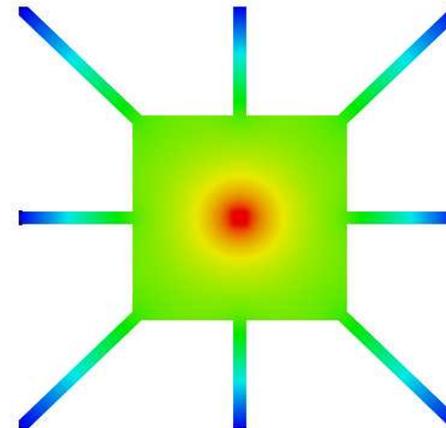
Model name: square\_100UM\_thermister  
Study name: shell2D  
Plot type: Thermal - Plot1  
Time step: 1 time: 0.001 Seconds



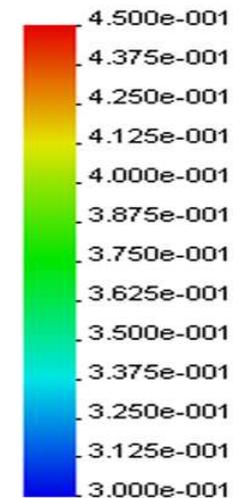
High G



Medium G

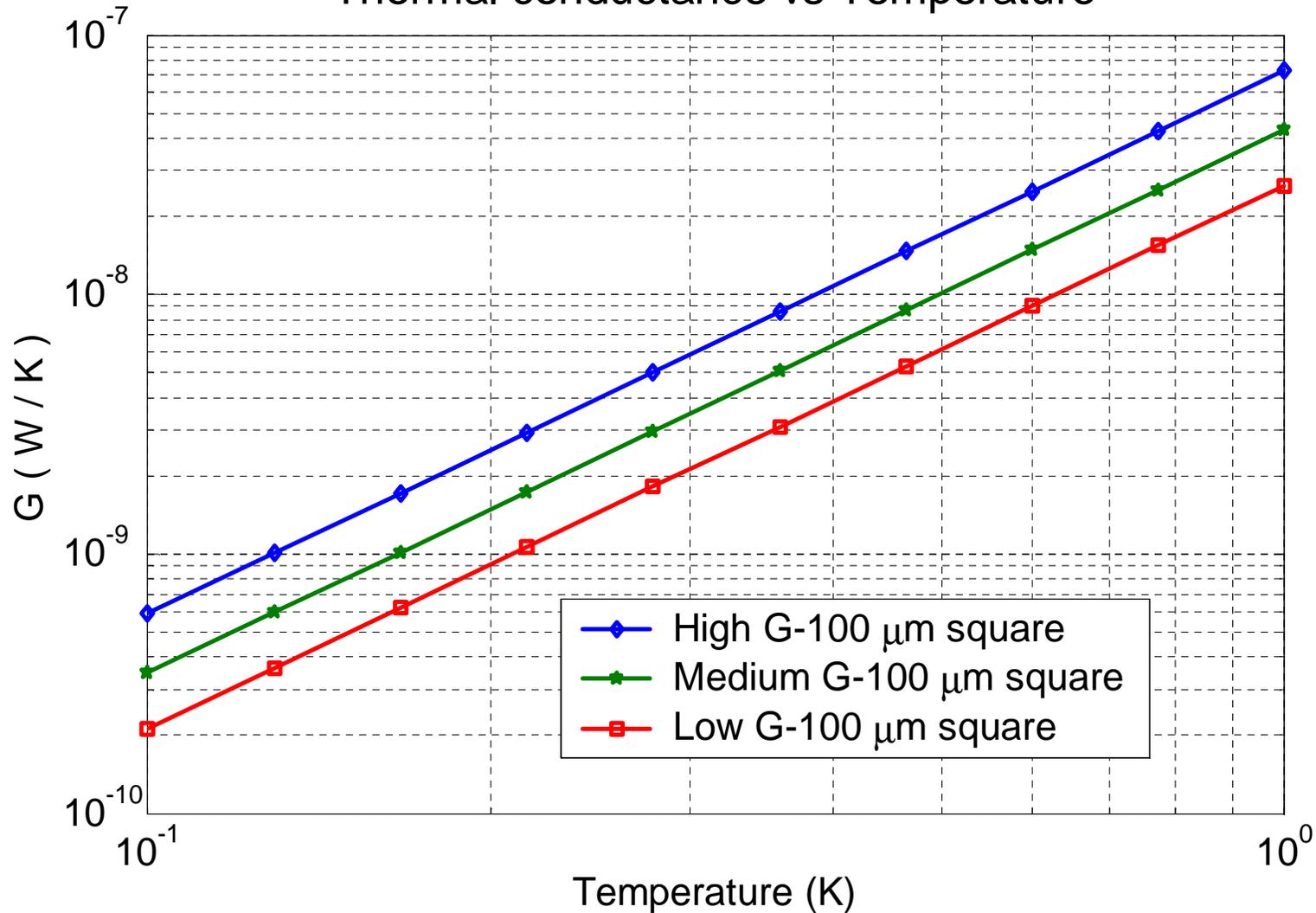


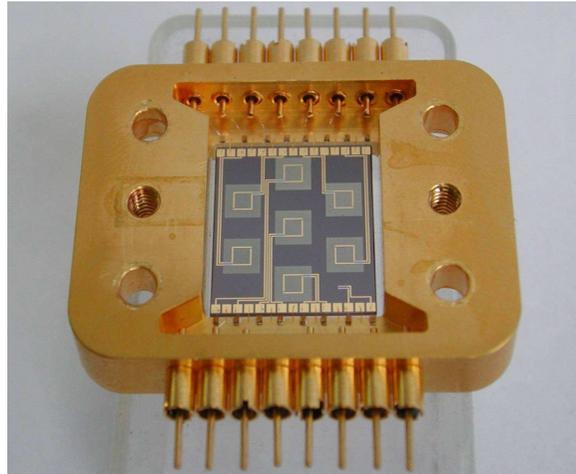
Low G



## Thermal conductance

Thermal conductance vs Temperature





7-element array



39-element array

Small Bolometer  
CAmera (SABOCA)

## APEX

## The Atacama Pathfinder Experiment

Max-Planck-Institut  
für  
Radioastronomie



12m - Submillimeter telescope APEX

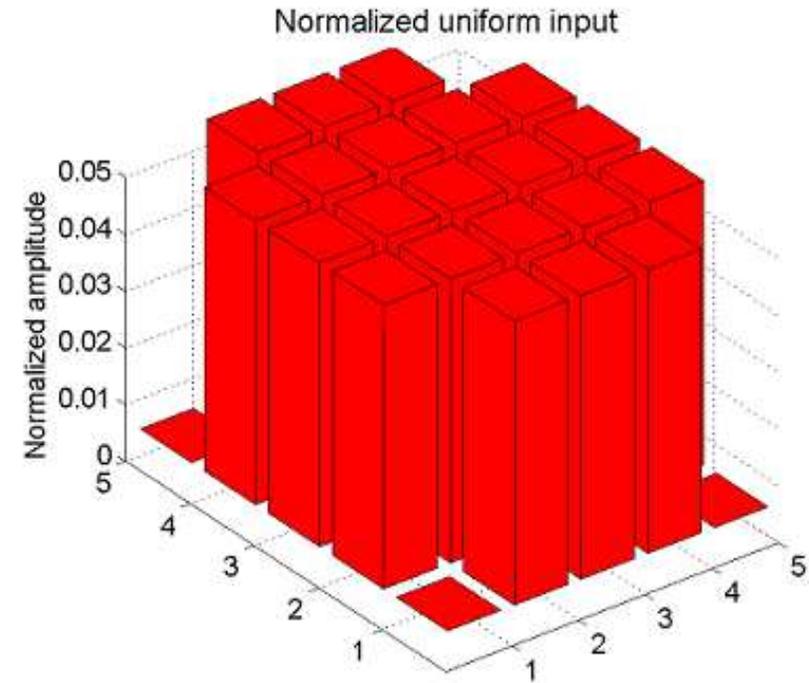
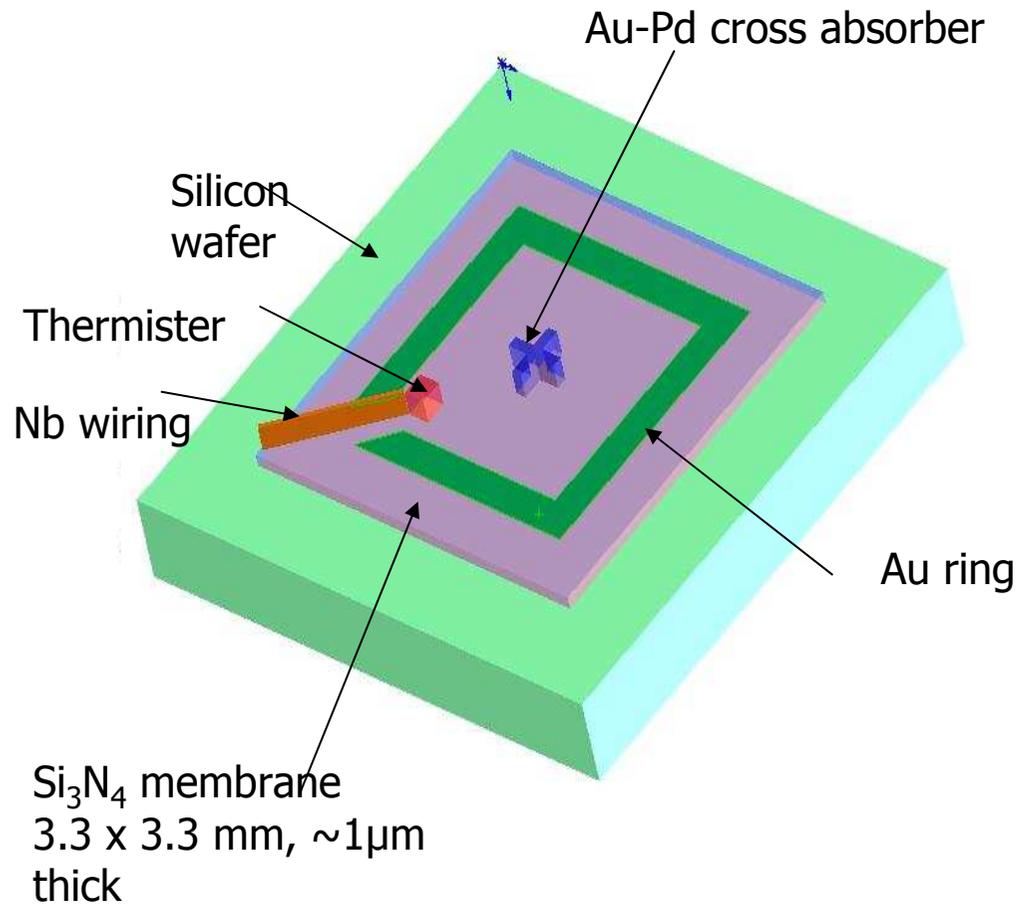
# Application: Astronomical detectors



## Chajnantor Plateau Atacama, Chile

Latitude            23 deg south  
Longitude         68 deg west  
Altitude            5100 m=16700 feet  
230 GHz → 1.5THz

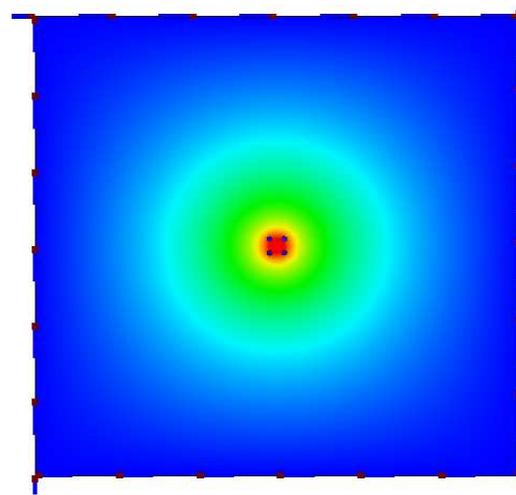




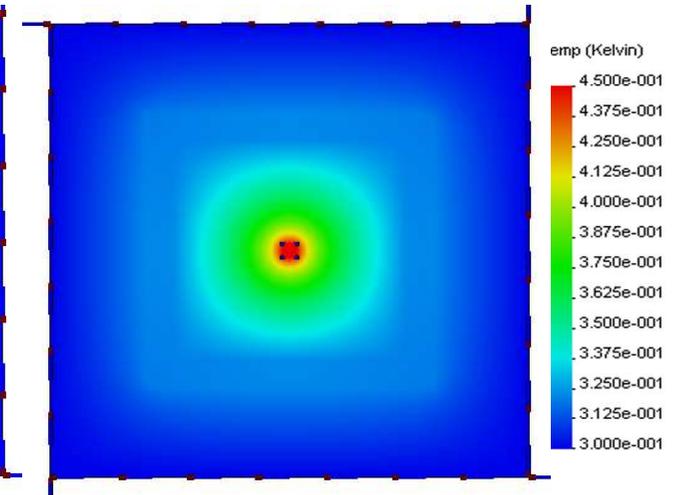
- FEA simulations showed increase in thermalization by depositing the center ring around radiation absorbing area.

- Detectors successfully used for Terahertz application.

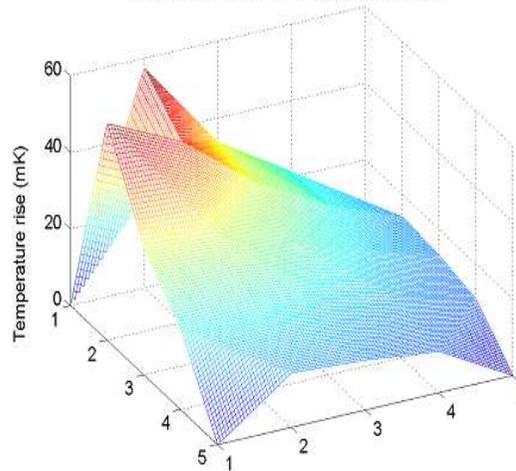
Without center Au ring



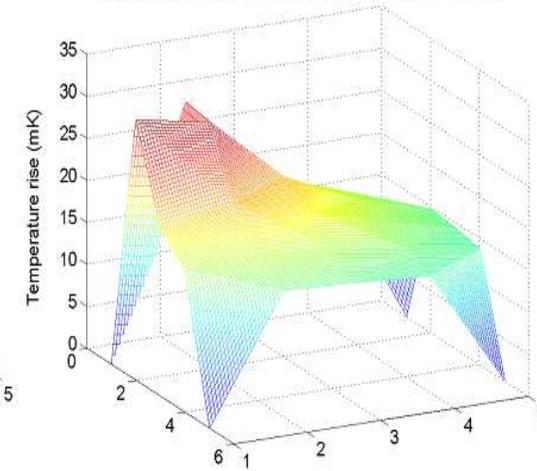
With center Au ring



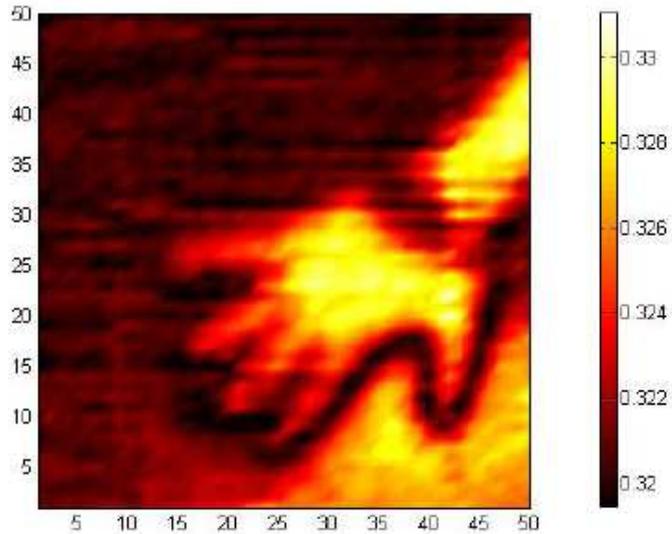
Surface fit to temperature distribution-High G



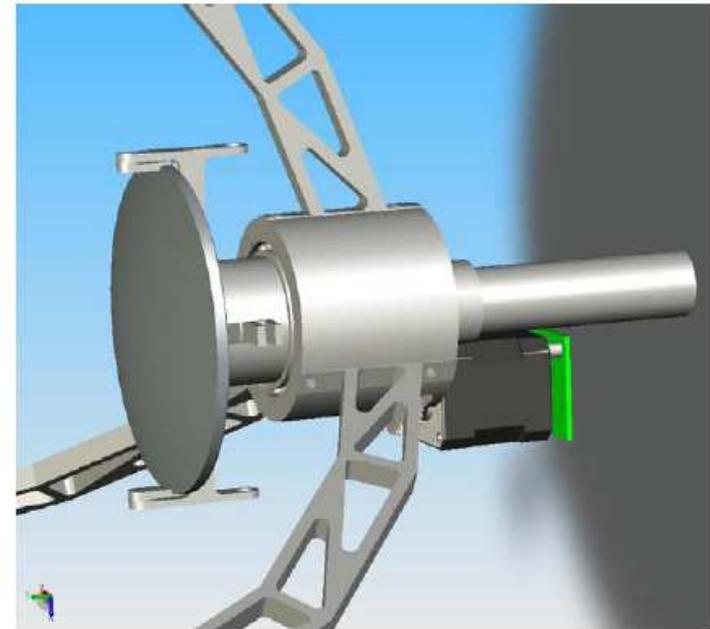
Surface fit to temperature distribution-High G-Au ring



# Application: THz scanner

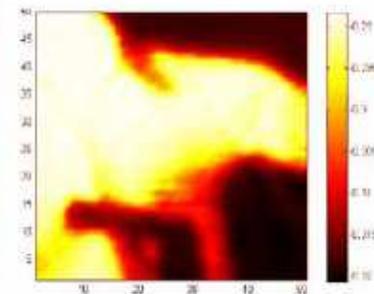


Design of an optical scanner based on a gyrating mirror



Human hand at 0.34 THz, from 5 meter distance

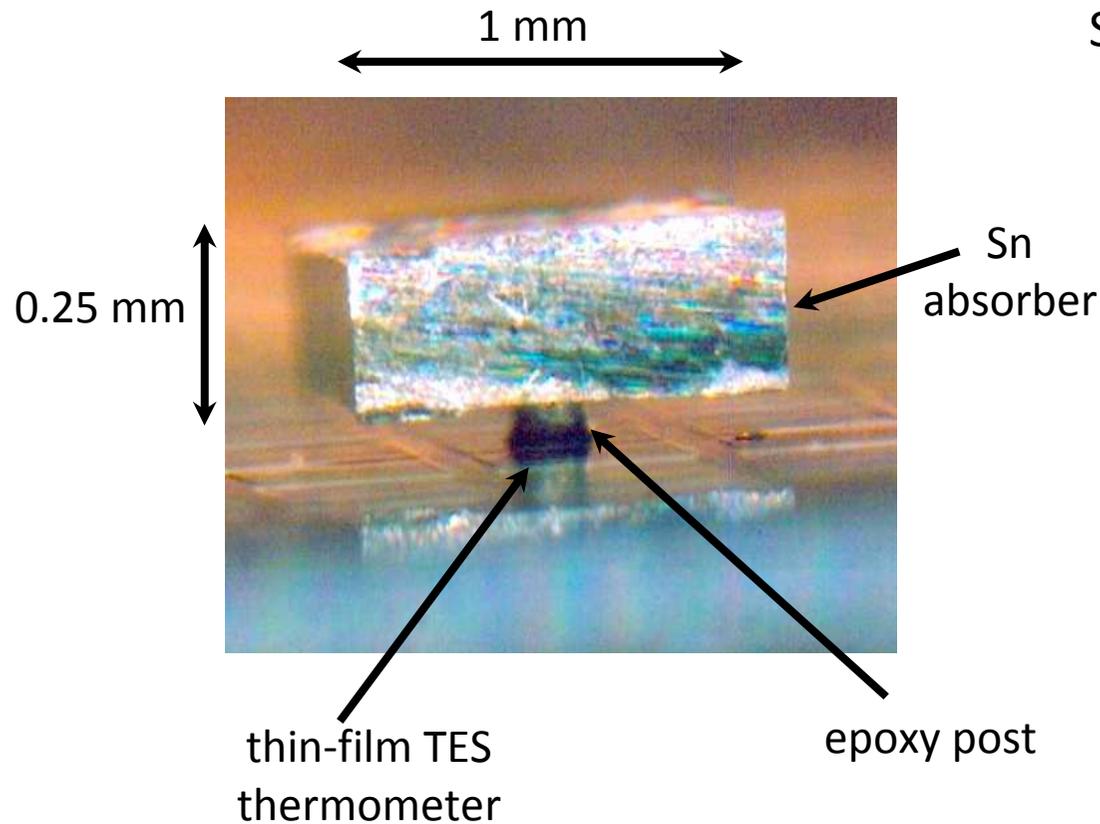
Test scenario: a test person hides a handgun mock-up underneath its shirt, and the THz scan reveals the threat from 5 meter distance.



To stop gamma-rays, a bulk absorber is required

Microcalorimeters can provide  $\Delta E < 50$  eV FWHM at 100 keV !

State-of-art germanium = 500 eV

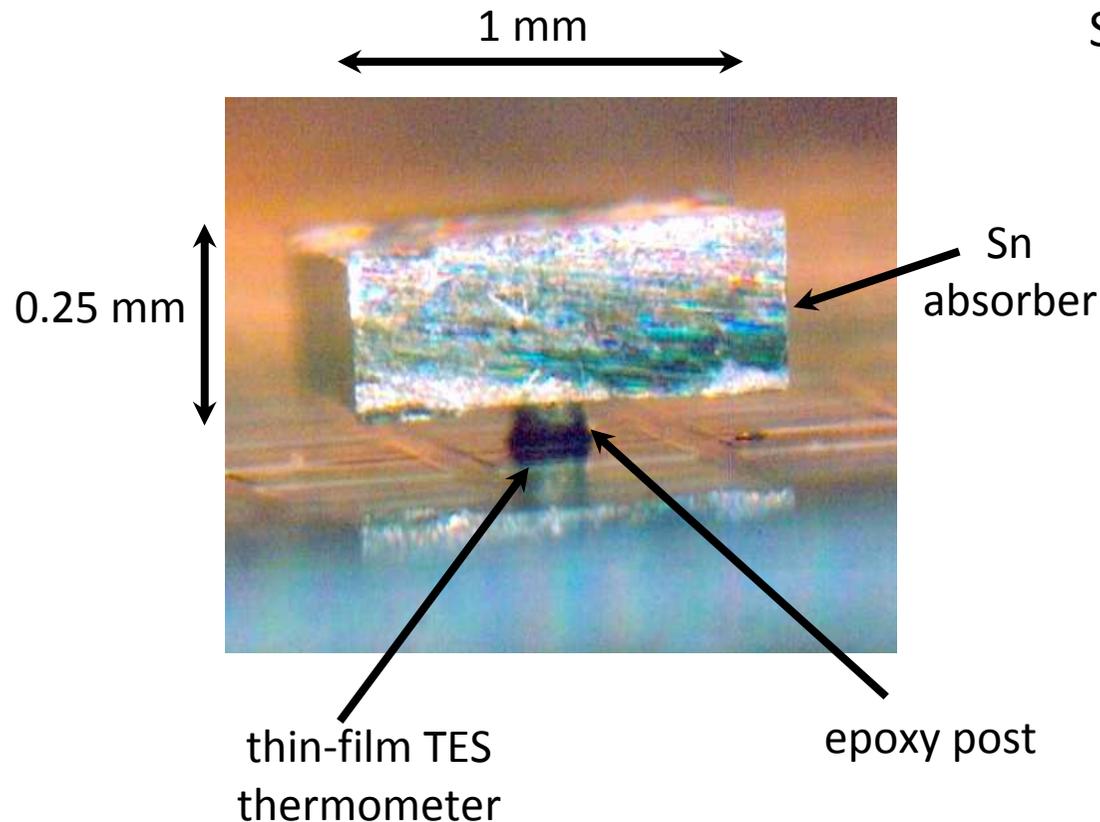


	$\mu$ cal	Ge	
$\Delta E$	50 eV	500 eV	
area	1-2 mm <sup>2</sup>	~500 mm <sup>2</sup>	
$\Gamma$	~100 Hz	~50 kHz	
QE	25% (at 100 keV)	~95%	

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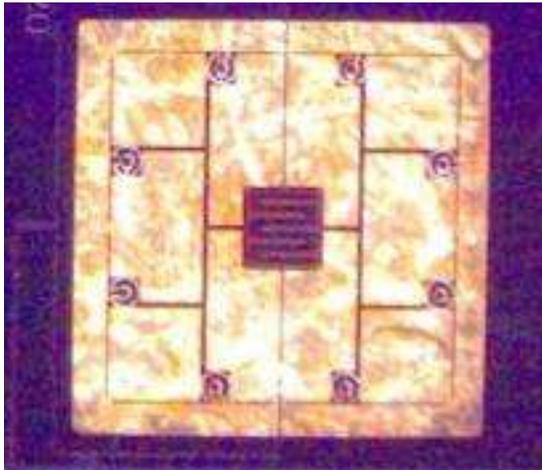
State-of-art germanium = 500 eV



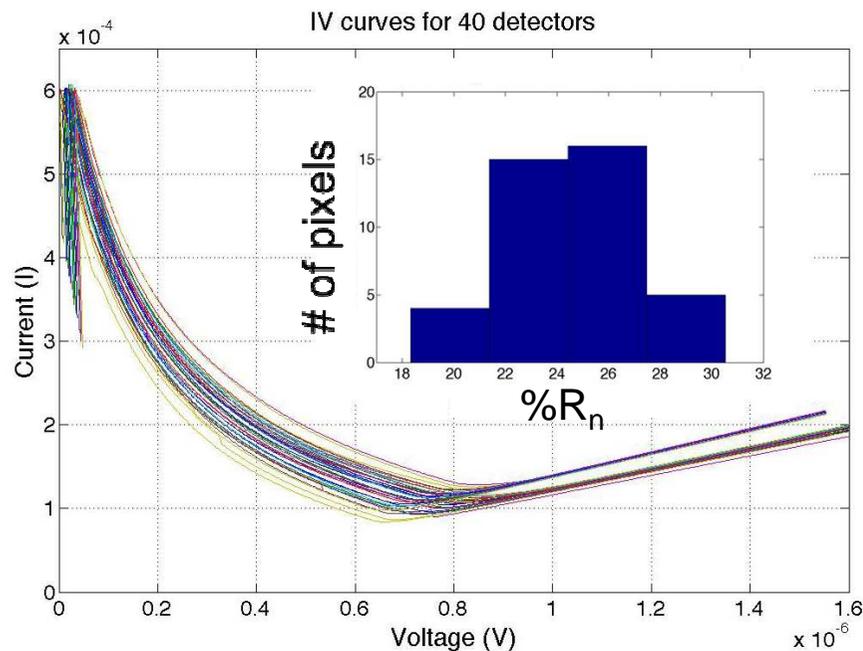
	$\mu$ cal	Ge	256 ucal
$\Delta E$	50 eV	500 eV	50 eV
area	1-2 mm <sup>2</sup>	~500 mm <sup>2</sup>	~500 mm <sup>2</sup>
$\Gamma$	~100 Hz	~50 kHz	~25 kHz
QE	25% (at 100 keV)	~95%	25%

Microcalorimeter arrays may perform better than Ge for some apps – Pu isotopics

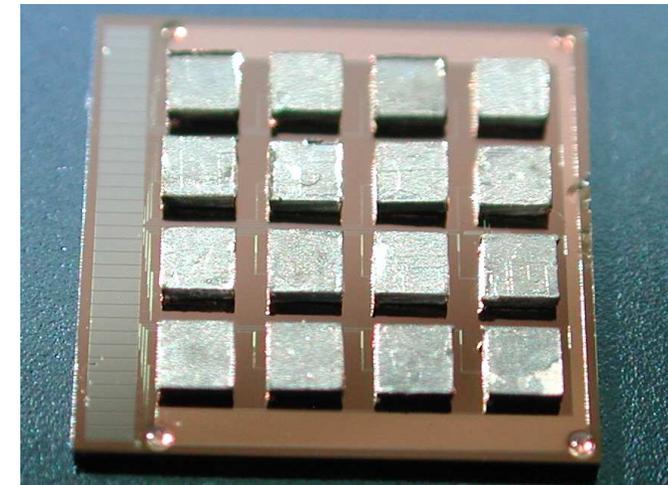
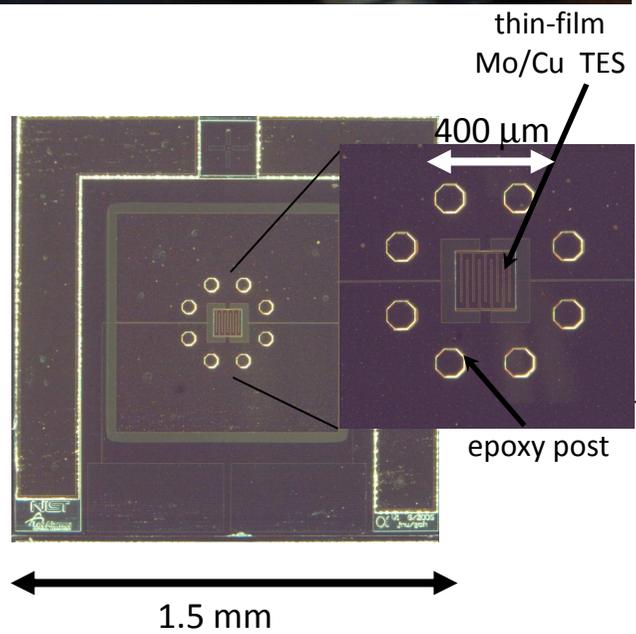
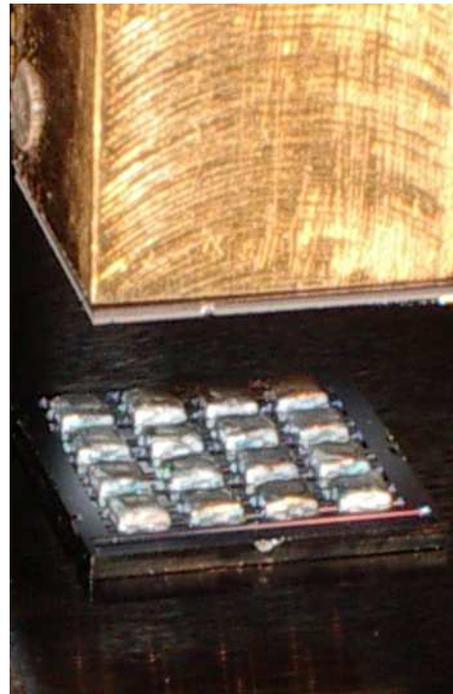
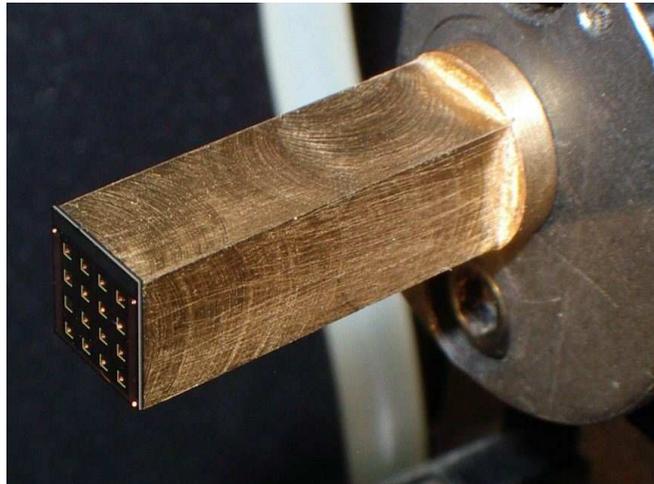
Microcalorimeter arrays may enable new applications – Pu in spent reactor fuel



- multi-post design
  - better mechanical support for the absorber
  - reduced mechanical and chemical stress on the TES thermometer
  - reduced dependence of pulse height on interaction position
- larger absorber area for higher detection efficiency ( $2.25 \text{ mm}^2$  vs.  $1 \text{ mm}^2$ )
- larger dynamic range, better linearity

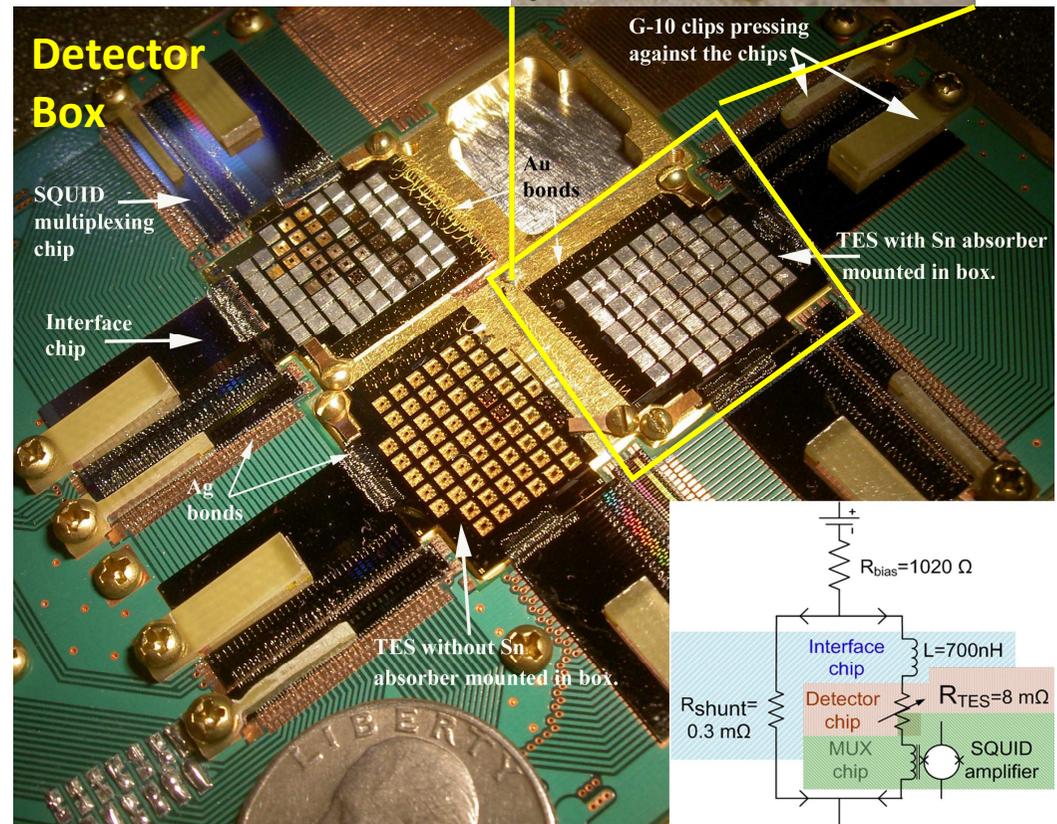
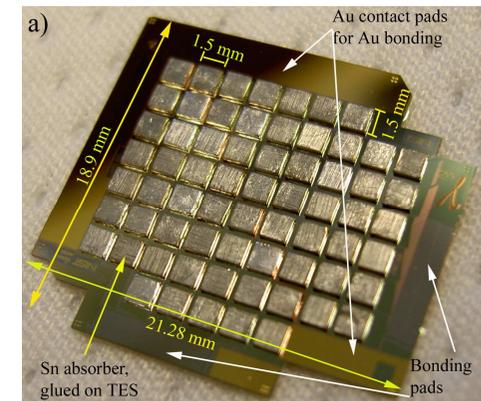


# Application: $\gamma$ -ray spectroscopy



## Gamma-ray detectors

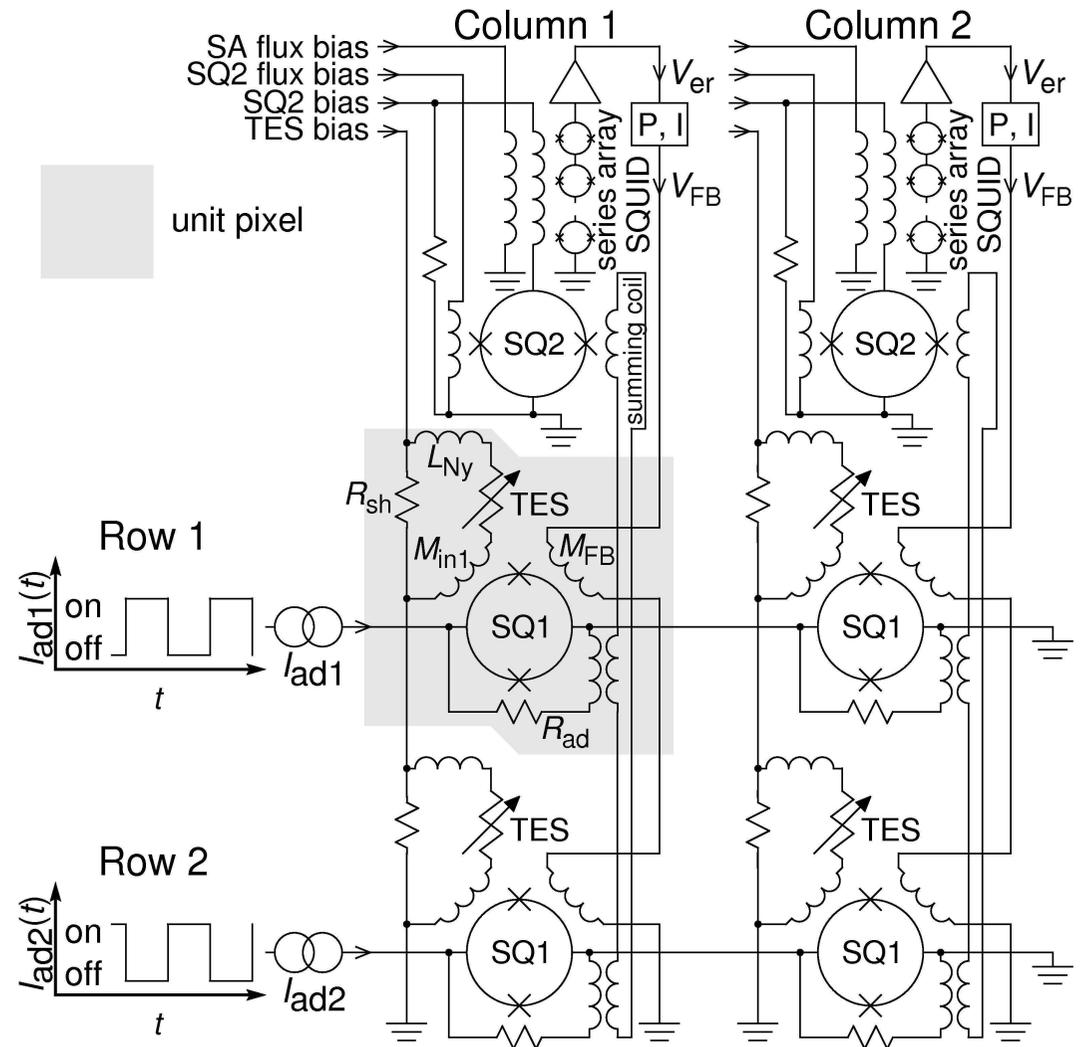
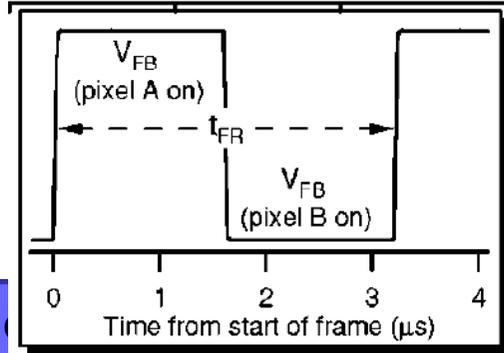
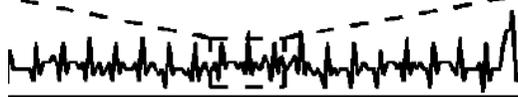
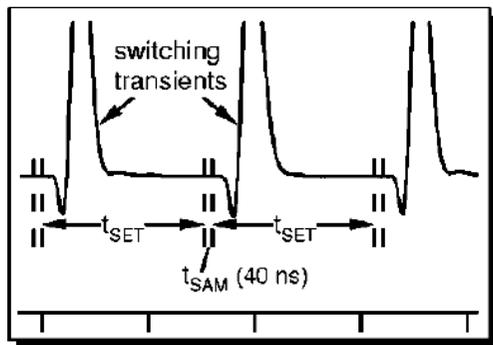
- 55 working detectors out of total 66 detectors.
- 31 Multiplexing detectors.
- Linear from 60 KeV – 130 KeV region.



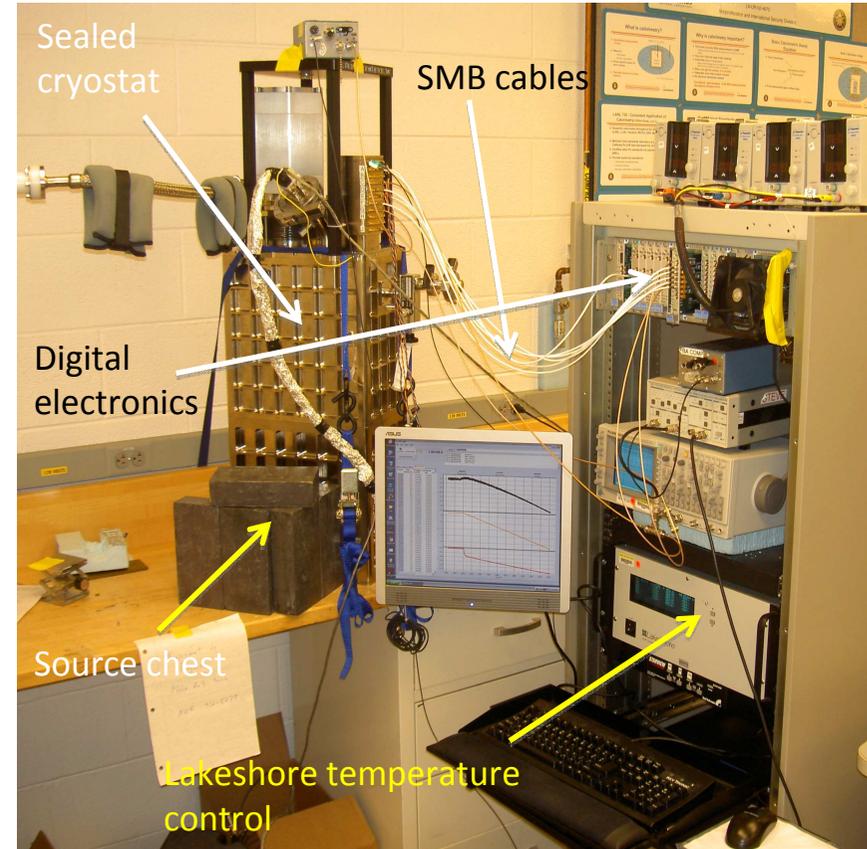
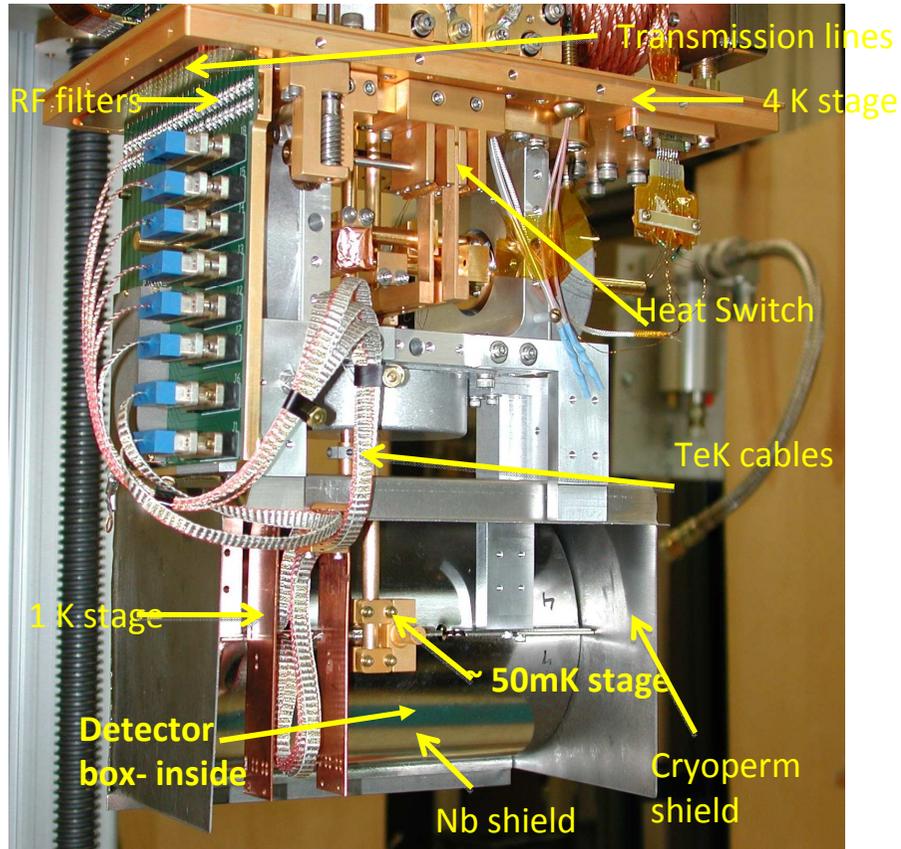
## Time division multiplexing: SQUIDs

- Three stage SQUID wiring.
- Flux-locked loop.

$$t_{FR} = N_{MUX}(t_{SET} + t_{SAM})$$

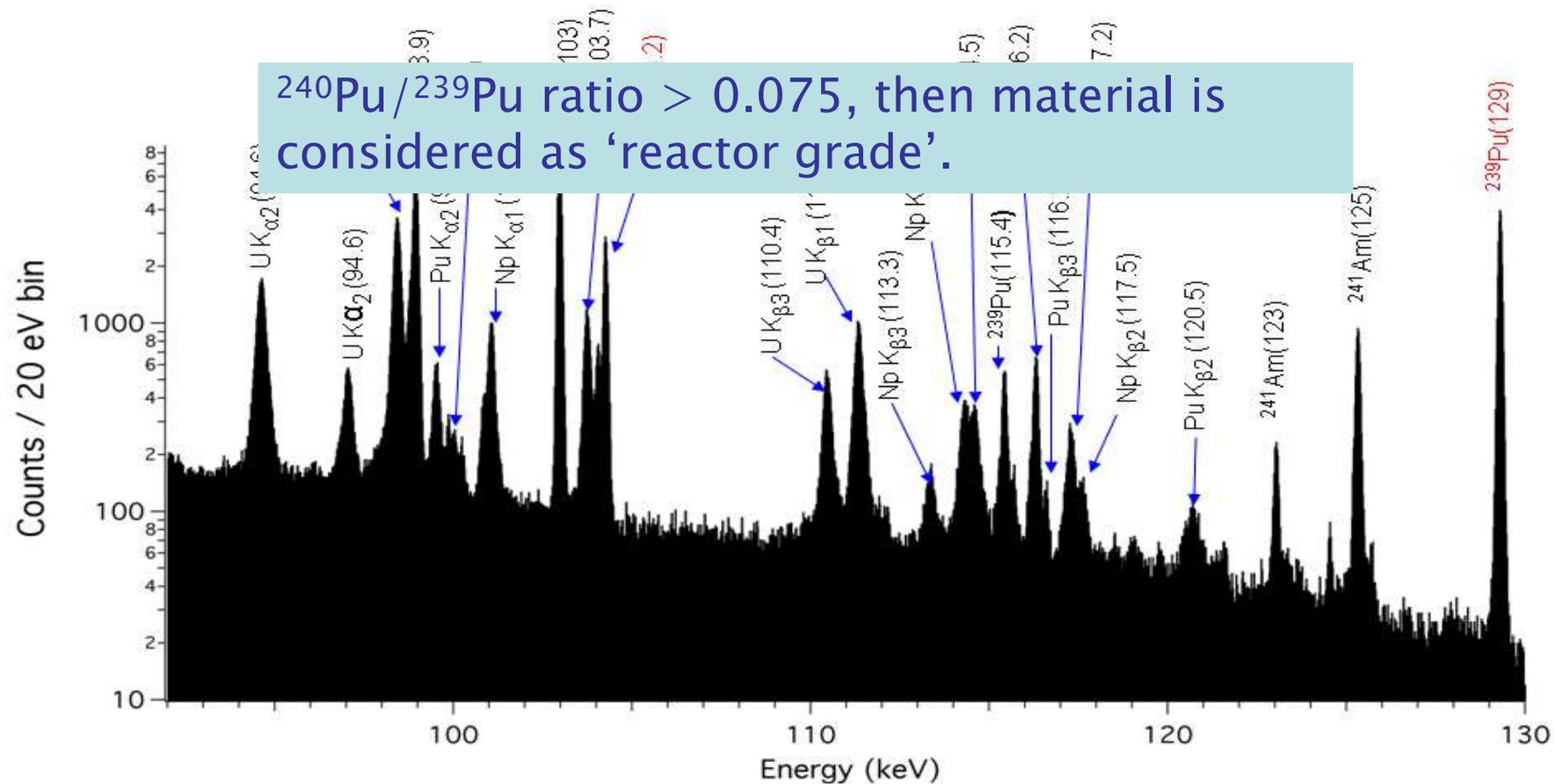


## Gamma-ray cryogenics



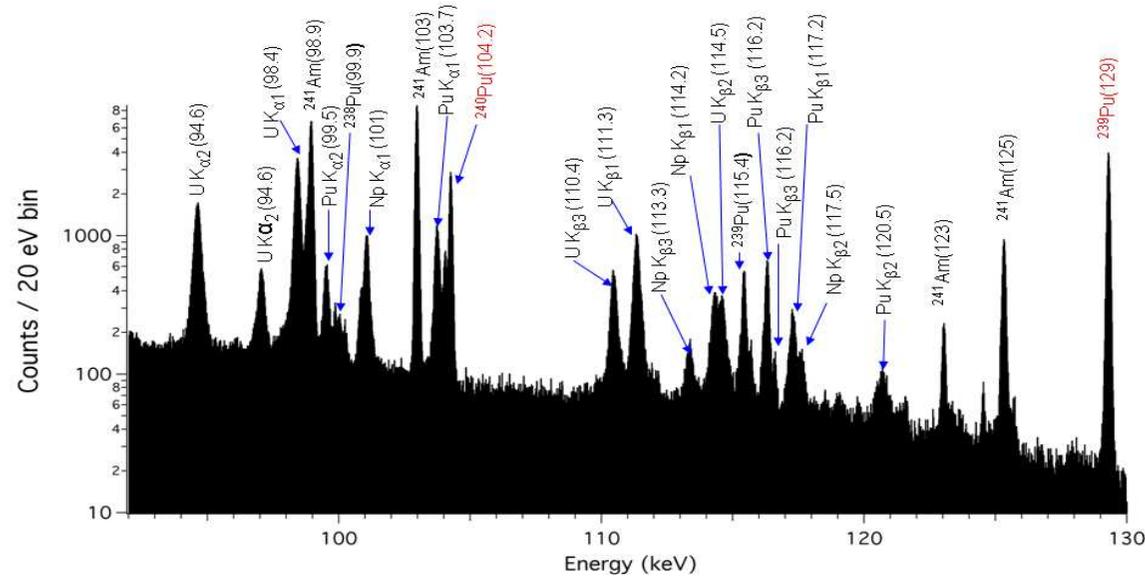
- Configuration allowing 264 detectors.
- Typical base temperature=62mK, Temperature regulated at 100mK for 20 hours.
- No liquid cryogenics; units working continuously at Los Alamos and NIST labs.

## Plutonium Spectrum:



- 1) Total: 38.9 million pulses. 2)  $\sim 18,700$   $^{240}\text{Pu}$  counts  $\rightarrow$  for the first time using TES detectors. 3)  $\Delta E \sim 120$  eV. Expected resolution: 70 eV

## Plutonium Spectrum:



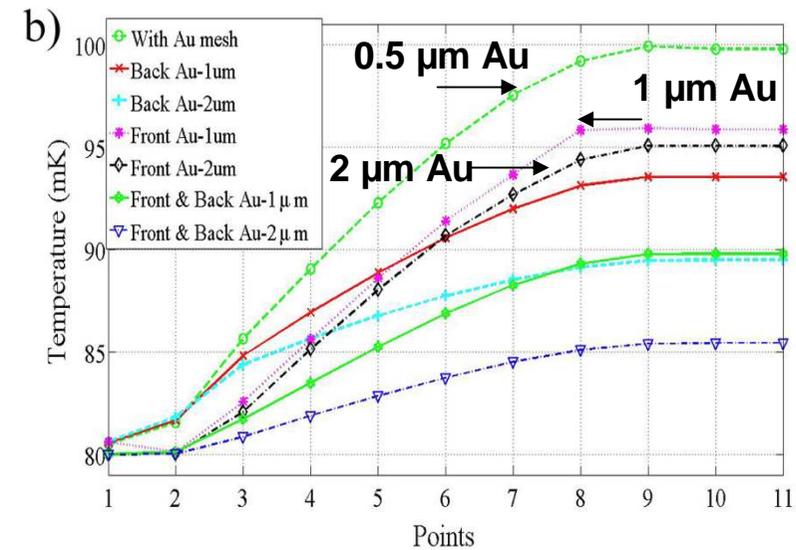
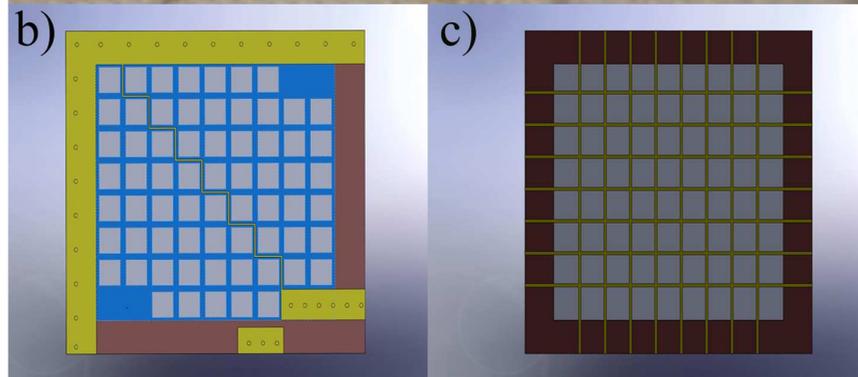
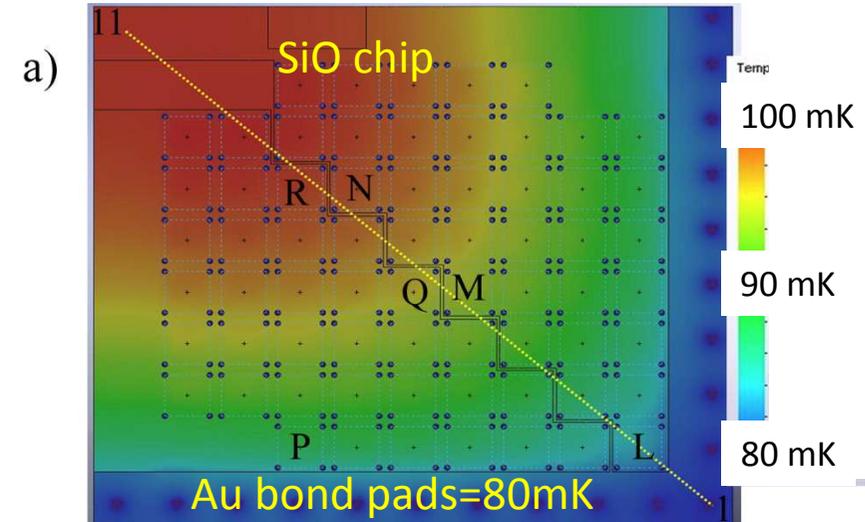
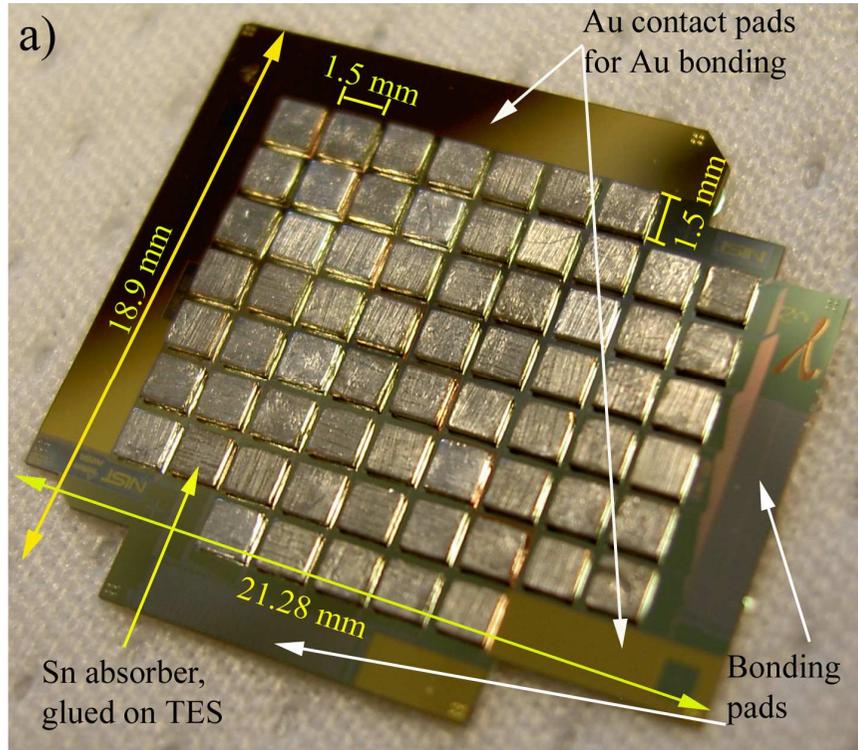
### Things to improve:

- Number of pixels (can be achieved by circuit inductance and bandwidth optimization).
- Energy Resolution.

Expected 70 eV, Achieved: 120 eV.

→ Thermal cross-talk between the pixels can degrade the energy resolution.

# Application: $\gamma$ -ray spectroscopy



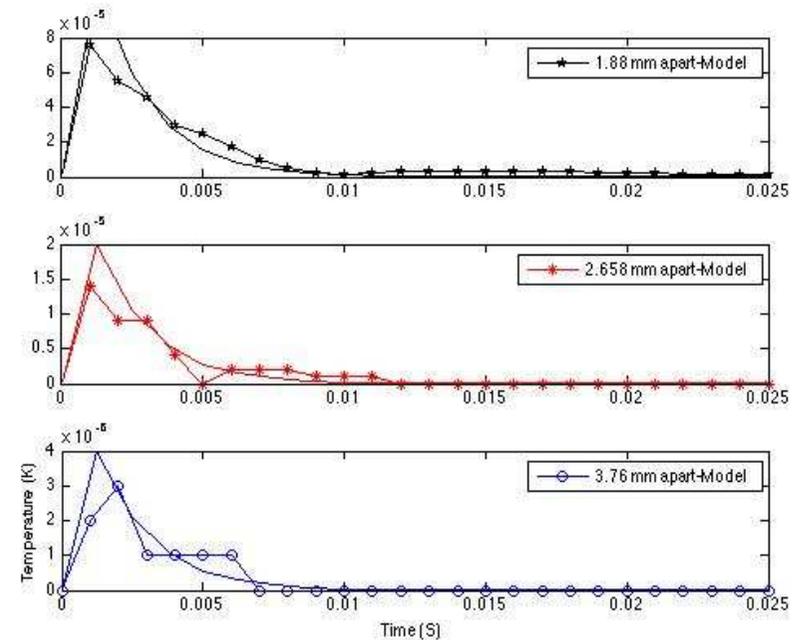
## Thermal Cross-talk determination

- 2% thermal cross-talk in neighboring pixel.
- Better thermalization of chip should increase the resolution by 8-10 %.
- New chip is fabricated and is being tested at LANL.

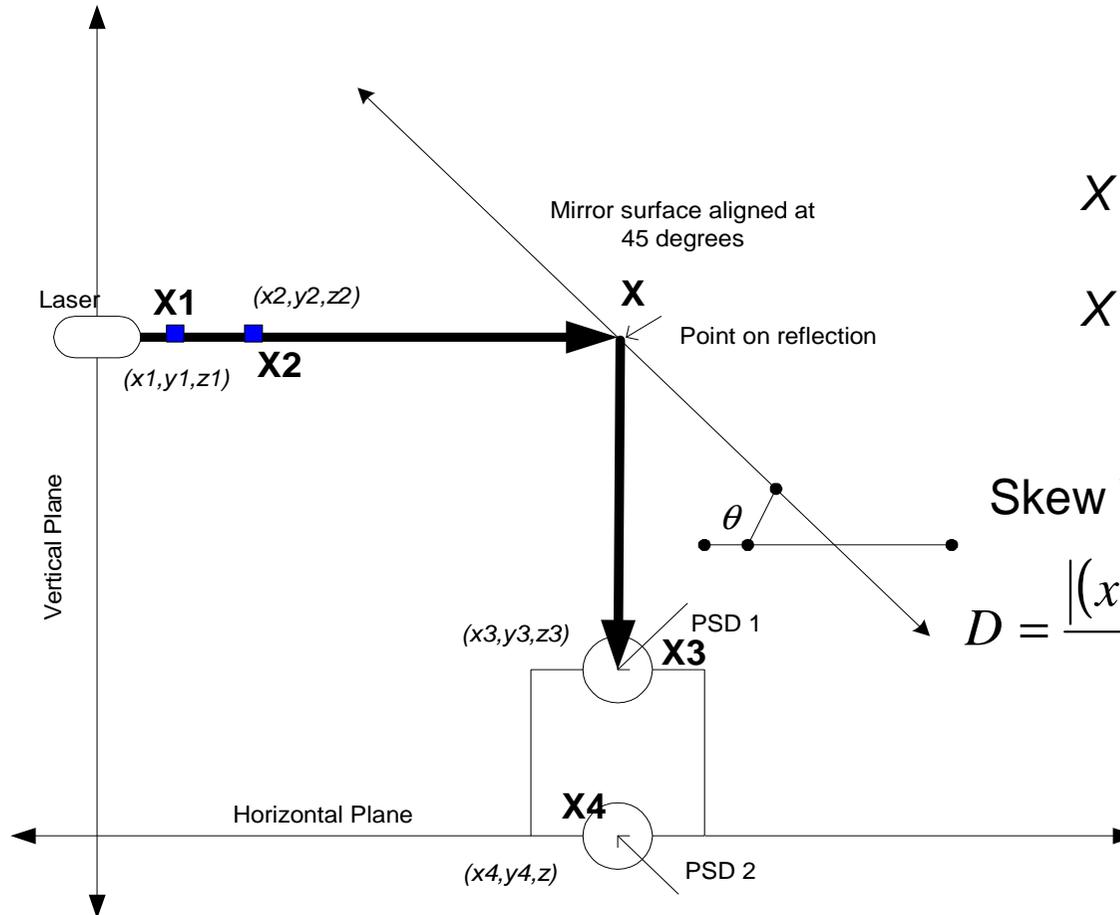
Current across TES

$$\text{(cross-talk)} = \text{IFFT} \left( T_{Si}(w) \cdot \frac{dT_{TES}}{dT_{SiO}} \cdot \frac{dI(w)}{dT(w)} \right)$$

$$T(\rho, t) = \int \underbrace{2 \cdot (e^{-t'/\tau_F} - e^{-t'/\tau_R})}_{\text{Input heat pulse}} \cdot \underbrace{\frac{E}{c} \cdot \frac{e^{(-\rho^2 / 4D(t-t'))} \cdot e^{-\gamma \cdot (t-t')}}{4\pi Dc(t-t')}}_{\text{Heat flow in Si}} \cdot dt'$$



## Non-Contact Surface Measurement



$$X = X1 + (X2 - X1) \times S$$

$$X = X3 + (X4 - X3) \times T$$

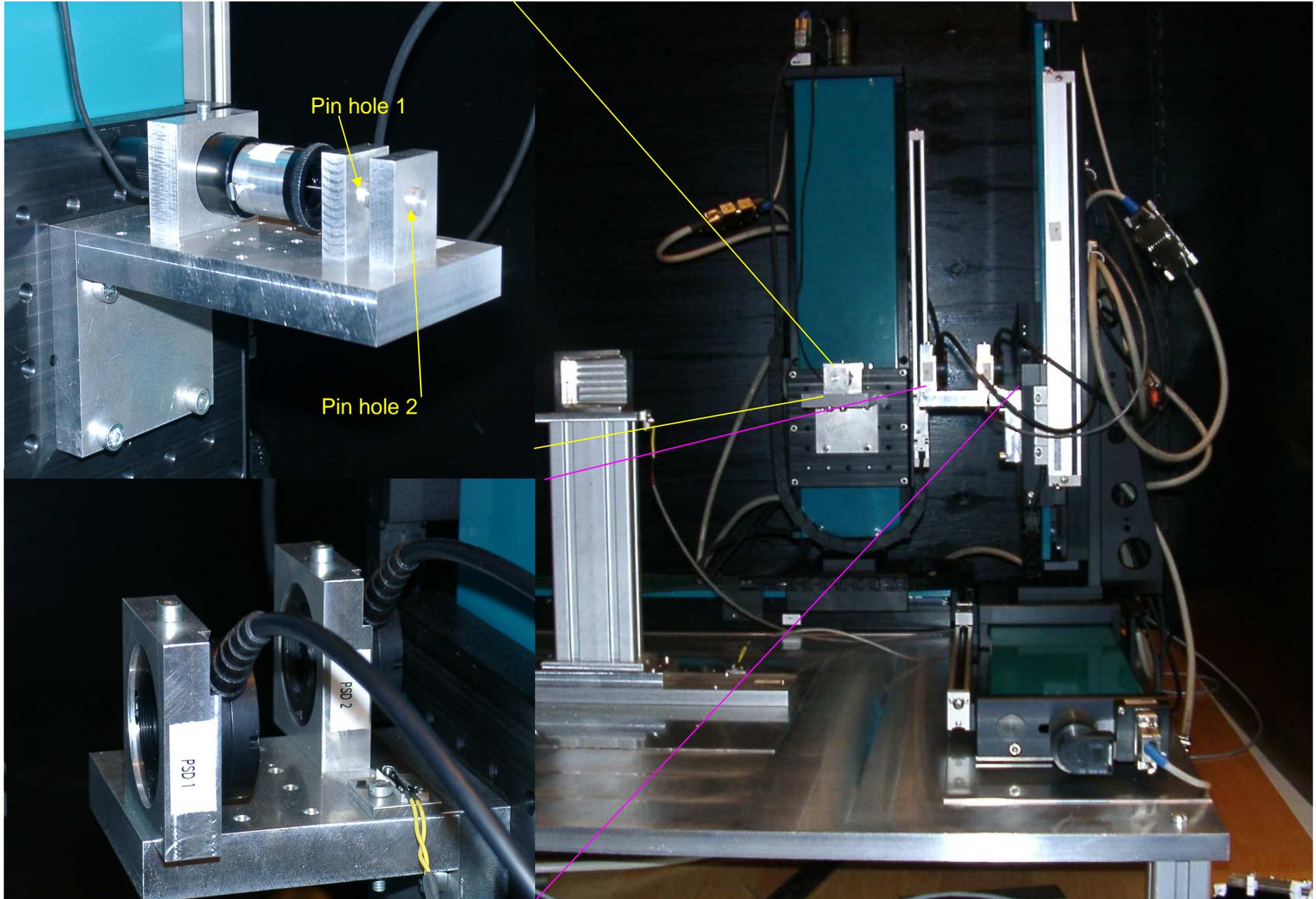
Skew Vector case:

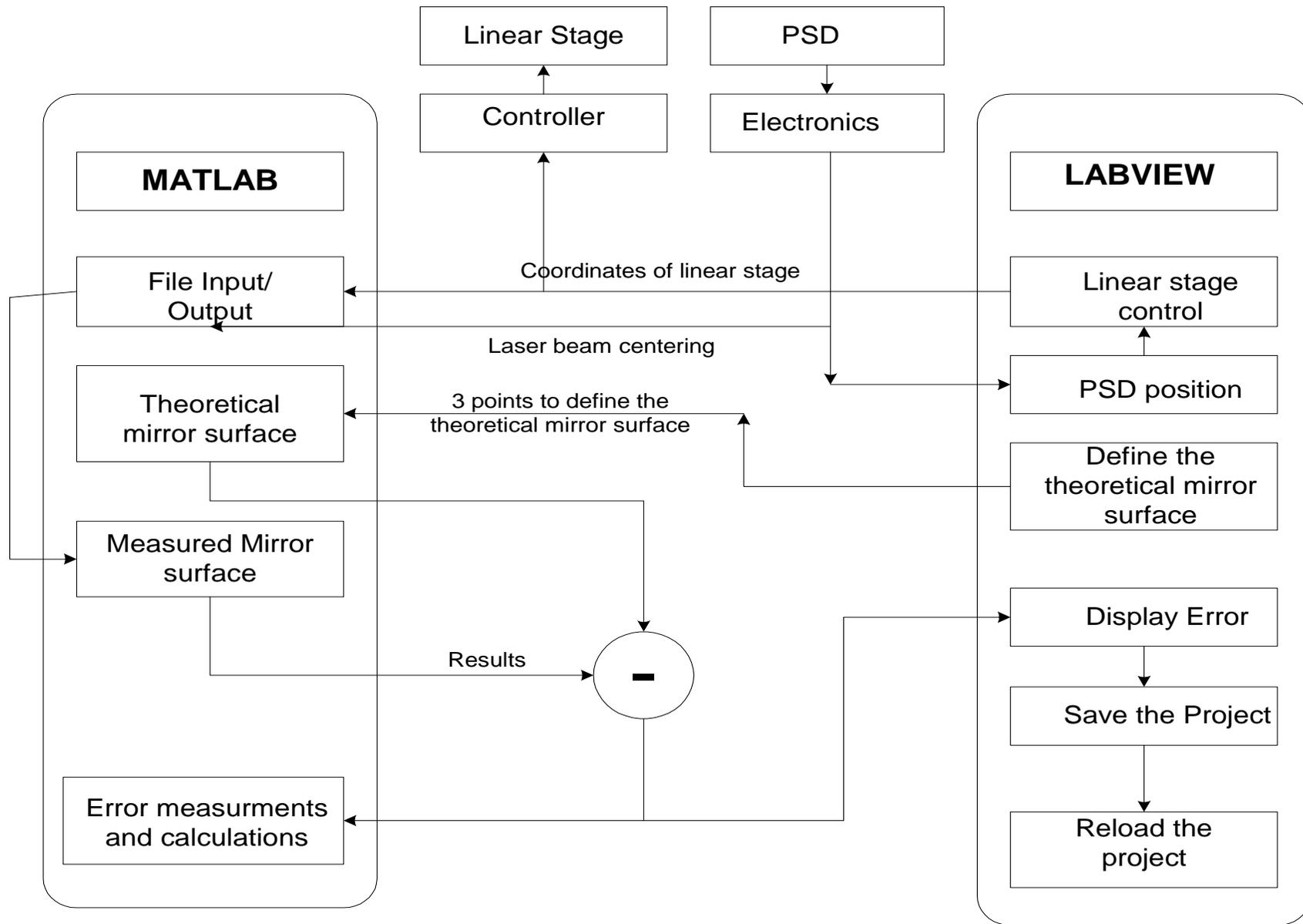
$$D = \frac{|(x1 - x3) \cdot [(x2 - x1) \times (x4 - x3)]|}{\|[(x2 - x1) \times (x4 - x3)]\|}$$

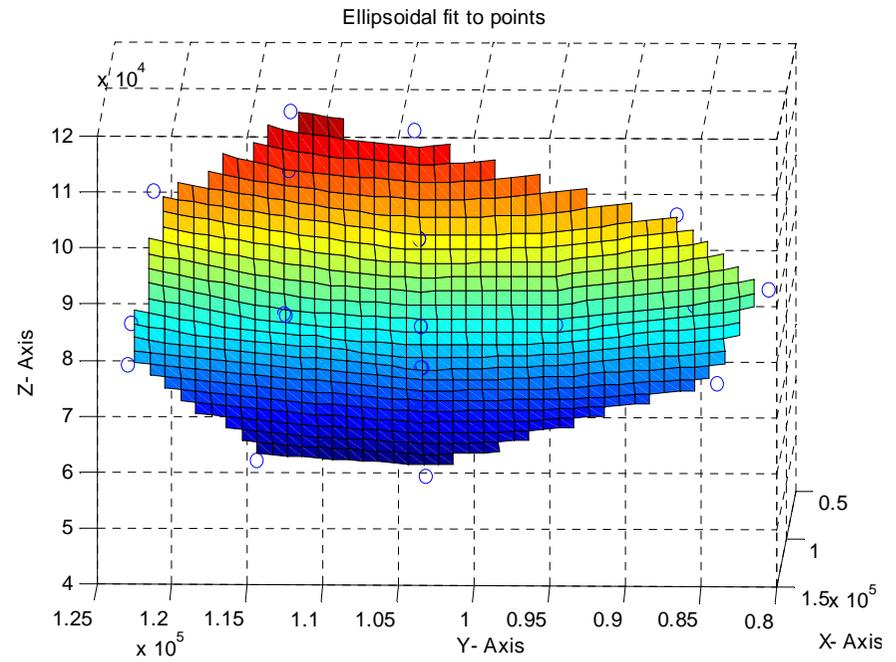
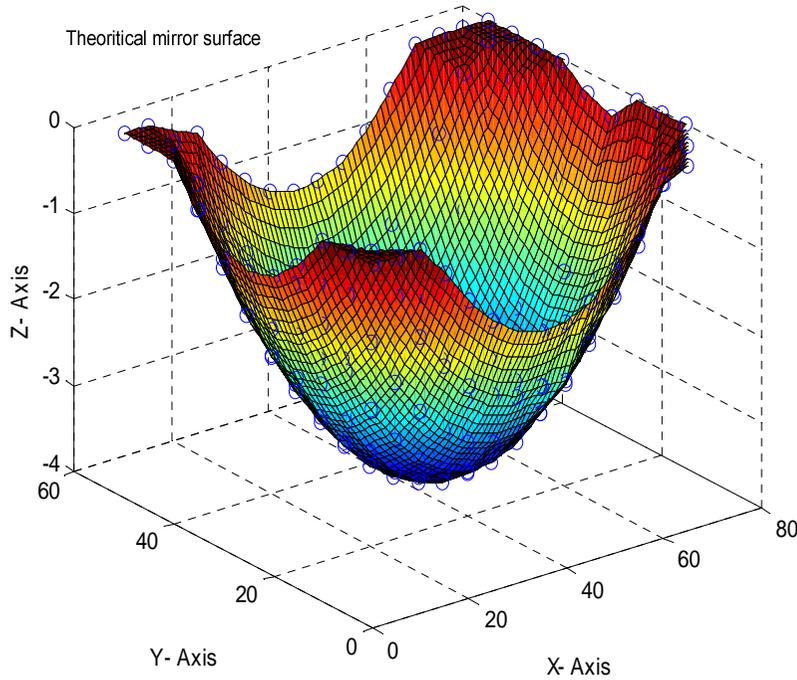
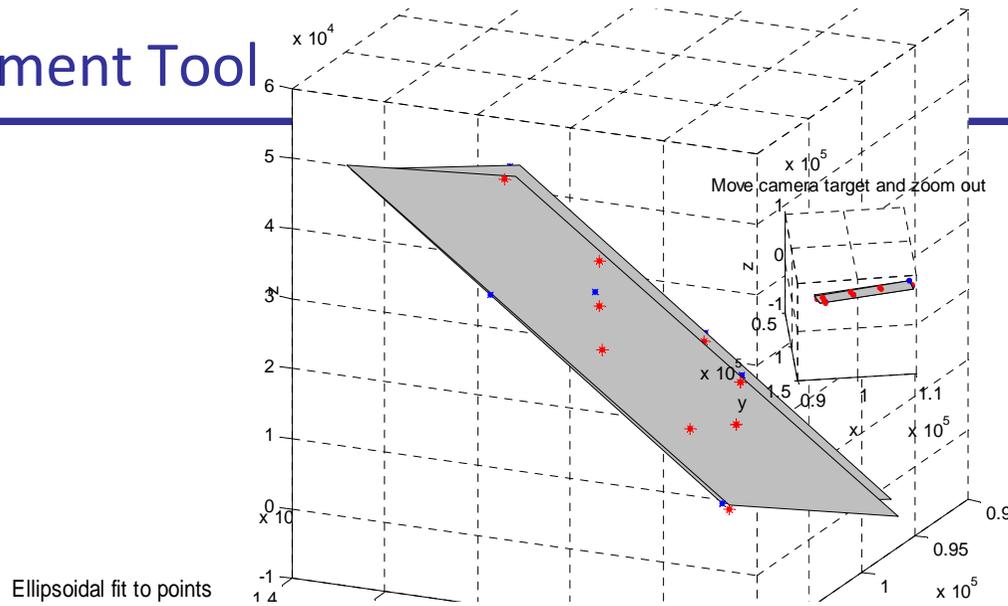
Perform this procedure to obtain 15 to 20 measured points on mirror surface

PSD: Position Sensitive Detector

# Optics Alignment Tool







**Theoretical curved mirror**

**Observed curved mirror**

## Conclusions:

- Superconducting TES are the most sensitive detectors for millimeter-submillimeter wavelength regime.
- Detectors successfully used for nuclear line forensics and spectroscopy → non-destructive analysis.

## Personal gain:

- Ability to work on a project as a team member and individual.
- Knowledge of software and hardware.
- Lab-work, Cryostats, cryogenics, instruments.

Everything has been very exciting and want to learn more.